

AD-A286 253

October 1994

A2

72

9.57.10-2.0

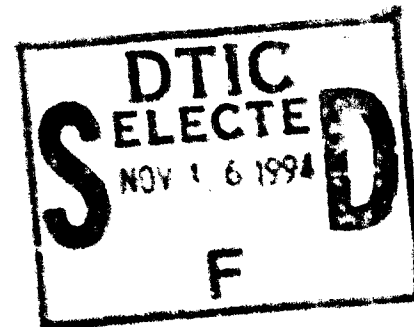
INVESTIGATION OF THE IONOSPHERIC
SHORT-TERM VARIABILITY

Final Technical Report

by

ZVI HOUNINER

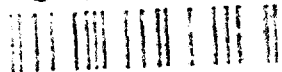
September 1993 - August 1994



United States Army
EUROPEAN RESEARCH OFFICE OF THE US ARMY
London, England
Contract number: DAJA-93-C-0035

Principal Investigator: Prof. G. Shviv
ASHER SPACE RESEARCH INSTITUTE
Technion, Haifa 32000, Israel

94-35316



Reproduced From
Best Available Copy

Approved for public release; distribution unlimited.

DTIC QUALITY INSPECTED 9

October 1994

AD

**INVESTIGATION OF THE IONOSPHERIC
SHORT-TERM VARIABILITY**

**Final Technical Report
by
ZWI HOUMINER**

September 1993 - August 1994

**United States Army
EUROPEAN RESEARCH OFFICE OF THE US ARMY
London, England
Contract number: DAJA-93-C-0035**

**Principal Investigator: Prof. G. Shaviv
ASHER SPACE RESEARCH INSTITUTE
Technion, Haifa 32000, Israel**

Approved for public release; distribution unlimited.

REPORT DOCUMENTATION PAGE			Form Approved OMB No. 0704-0188	
<small>Public reporting burden for this collection of information is estimated to average 1 hour per response, including the time for reviewing existing data sources, gathering existing data, reviewing the material for accuracy, completing and reviewing the collection of information, and sending the collection of information to the agency that collects the information. Send comments regarding this burden estimate or any aspect of this collection of information, including suggestions for reducing this burden, to Washington Headquarters Office, Directorate for Information Operations and Reports, 1215 Jefferson Davis Highway, Suite 1204, Arlington, VA 22202-4302, and to the Office of Management and Budget, Paperwork Project (0704-0188) in Washington, DC 20503.</small>				
1. AGENCY USE ONLY (Leave blank)	2. REPORT DATE October 1994	3. REPORT TYPE AND DATES COVERED Final Technical Report		
4. TITLE AND SUBTITLE Investigation of the Ionospheric Short-term Variability		5. FUNDING NUMBERS DAJA 45-913-C-0035		
6. AUTHOR(S) Z. Houminer				
7. PERFORMING ORGANIZATION NAME(S) AND ADDRESS(ES) Asher Space Research Institute, Technion, Haifa 32000, ISRAEL		8. PERFORMING ORGANIZATION REPORT NUMBER		
9. SPONSORING / MONITORING AGENCY NAME(S) AND ADDRESS(ES) European Research Office, U.S. Army 223 Old Marylebone Rd, London NW6 5TH		10. SPONSORING / MONITORING AGENCY REPORT NUMBER		
11. SUPPLEMENTARY NOTES Report no. taken from Form 50 R/D 7059-EE-01				
12a. DISTRIBUTION AVAILABILITY STATEMENT		12b. DISTRIBUTION CODE		
13. ABSTRACT (Maximum 200 words) See next page.				
14. SUBJECT TERMS		15. NUMBER OF PAGES		
		16. PRICE CODE		
17. SECURITY CLASSIFICATION OF REPORT	18. SECURITY CLASSIFICATION OF THIS PAGE	19. SECURITY CLASSIFICATION OF ABSTRACT	20. LIMITATION OF ABSTRACT	

ABSTRACT

Both the total electron content of the ionosphere (TEC) and the critical frequency of the F2 layer (foF2) exhibit large day-to-day variations during quiet and active geomagnetic periods. It is of great interest to ascertain whether good correlation exists between TEC daily variability about the monthly mean and foF2 variations. With the availability of the global GPS constellation to provide instantaneous time-delay values such a correlation may enable the improvement of HF short-term predictions using passive monitoring of TEC.

To determine the correlation, a pilot study was conducted using several months of TEC data taken in Haifa, Israel during 1980 as well as GPS time-delay measurements taken during the summer of 1992 in Jerusalem, Israel. The corresponding foF2 measurements were from Cape Zevgary, Cyprus.

The analysis showed, that for large percentages of the time very good correlation exists between TEC and foF2 short-term variations. The correlation coefficient varies from 0.7 - 0.8 during winter and summer months to about 0.5 - 0.6 during equinox months. A study of the diurnal dependence of the correlation indicates that better correlation exists during day-time than night-time. There was no indication that the correlation coefficient is dependent on geomagnetic activity during the period of this study.

The high correlation between TEC and foF2 in the limited data analysed, indicates, that real-time ionospheric HF prediction improvements are feasible when using transionospheric time-delay measurements.

KEY WORDS

Ionospheric variability; HF propagation; HF short-term predictions; Total electron content (TEC); foF2; Transionospheric time-delay; GPS.

TABLE OF CONTENT

1. Introduction	1
2. Experimental results	1
3. Conclusions	2
4. List of publications	3
Appendix A	4
Appendix B	9

Accession For	
NTIS CRA&I	<input checked="" type="checkbox"/>
DTIC TAB	<input type="checkbox"/>
Unannounced	<input type="checkbox"/>
Justification	
By	
Distribution /	
Availability Codes	
Dist	Avail and/or Special
A-1	

1. INTRODUCTION

To assess the possible improvement of HF short-term predictions from transionospheric time-delay measurements, it is of great interest to ascertain whether good correlation exists between TEC daily variability about the monthly mean and foF2 variability. To determine such correlation, a pilot study was conducted using several months of TEC data taken in Haifa, Israel during 1980 as well as GPS time-delay measurements taken during the summer of 1992 in Jerusalem, Israel. The corresponding foF2 measurements were from Cape Zevgari, Cyprus.

The TEC at Haifa was determined from Faraday rotation observations using the signal of the Sirio satellite. The geographic subionospheric point corresponding to a mean ionospheric height of 350 km (along the signal path) is 30.3°N and 28.9°E . The geographic coordinates of Cape Zevgari are 34.6°N and 32.9°E . The measurements are thus separated by 4.3° in latitude and 4° in longitude. The GPS observations were taken at the National Physical Laboratory in Jerusalem using a Ionospheric Measurement System developed at the National Institute of Standards and Technology (NIST) at Boulder, Colorado for accurate time transfer. This system provides 15 minutes averages of time-delay data to all GPS satellites above the horizon. The equivalent TEC from the time-delay measurements was determined only for satellites at elevations larger than 30° and having a subionospheric point along the line of sight within 5° of the latitude of the ionosonde. The observed time-delays were then corrected for satellite biases converted into vertical TEC and averaged to obtain hourly values.

2. THE EXPERIMENTAL RESULTS

For each month analyzed, the variability of both the TEC and foF2 was determined. The variability is calculated by subtracting the monthly average value from each hourly value and deviding by

the monthly average value. Using the hourly values of the variability, a cross-correlation analysis between the TEC and foF2 was performed as well as auto-correlation for both TEC and foF2. Finally, the variability of the slab-thickness τ , determined from the ratio $TEC/(foF2)^2$, was calculated.

The full account of the data analysis and the results obtained are summarised in the paper "Assessment of foF2 Short-term Variations from Transionospheric Time-delay Measurements" to be published in the Proceedings of the International Beacon Satellite Symposium, which took place at Aberystwyth, Wales, from 11-15 July 1994. The preprint of the paper is given in Appendix A.

A complete set of figures for each month analyzed in this study are given in Appendix B. The following figures are shown for every month:

- a. Hourly values of the variability of foF2 and TEC.
- b. Cross-correlation function between hourly values of foF2 and TEC.
- c. Auto-correlation functions for foF2 and TEC.
- d. Hourly values of the variability of the slab-thickness τ .

For the period June-August 1992 the results of smoothing the hourly values of foF2 and TEC by a 3-hour and 5-hour running mean, are also shown. The smoothing remarkably improved the correlation between foF2 and TEC, which is discussed in the paper in appendix A.

3. CONCLUSION

The high cross-correlation (>0.75) coefficient for foF2 and TEC in the limited data presented here raises the possibility that real-time TEC measurements may be used to update foF2 value determinations. The cross-correlation may even be higher if the geographic subionospheric point of the TEC measurement is closer

to the geographic point of the foF2 measurement, which introduces an error, in addition to the possible inherent measurements uncertainties. TEC measurements utilizing satellite emitted signals are passive in nature and do not burden the electromagnetic spectrum. In addition, the availability of the global GPS constellation to provide instantaneous time-delay, or equivalently TEC, values could provide an instantaneous updating of foF2 models on a global basis as well as on a regional basis. Such capability is important for HF communication along short, medium and long range paths.

4. LIST OF PUBLICATIONS

Z. Houminer & H. Soicher, "Improvement of foF2 short-term predictions from transionospheric time-delay measurements." National Radio Science Meeting, University of Colorado, Boulder, 5-8 January 1994.

H. Soicher & Z. Houminer, "Real-time ionospheric HF prediction improvements by passive means." 19th Army Science Conference, Orlando, Florida, 20-23 June 1994.

Z. Houminer & H. Soicher, "Assessment of foF2 short-term variations from GPS time-delay measurements." International Beacon Satellite Symposium, Aberystwyth, Wales, 11-15 July 1994.

APPENDIX A

"Assessment of foF2 short-term variations from transionospheric time-delay measurements".

Proceedings of the International Beacon Satellite Symposium,
September 1994.

ASSESSMENT OF f_oF_2 SHORT-TERM VARIATIONS FROM TRANSIONOSPHERIC TIME-DELAY MEASUREMENTS

Z. Houminer¹ and H. Soicher²

1. Asher Space Research Institute, Technion - Israel Institute of Technology,
Haifa 32000, Israel

Fax: +972 4 230956, e-mail: aszwih@vmsa.technion.ac.il

2. Space and Terrestrial Communications Directorate,
US Army Communications-Electronics Command, Fort Monmouth,
NJ 07703-5203, USA

Fax 1(908)532-0456, e-mail: Soicher@DOIM6@Monmouth - emh3.army.mil

1. INTRODUCTION

To assess the possible improvement of HF short-term predictions from transionospheric time-delay measurements, it is of great interest to ascertain whether good correlation exists between TEC daily variability about the monthly mean and f_oF_2 variability. To determine such correlation, a pilot study was conducted using several months of TEC data taken in Haifa, Israel during 1980 (Soicher et al., 1982) as well as GPS time-delay measurements taken during the summer of 1992 in Jerusalem, Israel. The corresponding f_oF_2 measurements were from Cape Zevgari, Cyprus.

The TEC at Haifa was determined from Faraday rotation observations using the signal of the Sirio satellite. The geographic subionospheric point corresponding to a mean ionospheric height of 350km (along the signal path) is 30.3°N and 28.9°E. The geographic coordinates of Cape Zevgari are 34.6°N and 32.9°E. The measurements are thus separated by 4.3° in latitude and 4° in longitude. The GPS observations were taken at the National Physical Laboratory in Jerusalem using a Ionospheric Measurement System developed at the National Institute of Standards and Technology (NIST) at Boulder, Colorado for accurate time transfer. This system provides 15 minutes averages of time-delay data to all GPS satellites above the horizon. The equivalent TEC from the time-delay measurements was determined only for satellites at elevations larger than 30° and having a subionospheric point along the line of sight within 5° of the latitude and longitude of the ionosonde. The observed time delays were then corrected for satellite biases using the JPL table of corrections (Wilson and Mannucci, 1993), converted into vertical TEC (Klobuchar, 1987) and averaged to obtain hourly values.

2. EXPERIMENTAL RESULTS

Figure 1 shows the hourly values of the variability of f_oF_2 and TEC for the period 1-16 August 1992. That period is near the maximum of solar cycle 22 (monthly mean sunspot number 103). The variability is determined by subtracting the monthly average value from each hourly value and dividing by the monthly average value. The results of cross correlation analysis on the f_oF_2 and TEC depicted in Figure 1 are shown in Figure 2. A maximum cross-correlation coefficient of 0.7 occurs at zero time lag while the correlation reduces very quickly with time lag. It is thus shown that the correlation between f_oF_2 and TEC is very good. Auto-correlation results for both f_oF_2 and TEC for the same time period (Figure 3) shows that the auto-correlation is similar in character to the cross correlation function and drops to 0.5 after a time lag of 3-4 hours. This is consistent with results obtained elsewhere indicating that short-term predictions based on real-time observations can only be useful for a few hours.

It can be seen from Figure 1 that both the variations in f_oF_2 and TEC are rather noisy on an hour-to-hour basis which may be caused by measurements and data reduction errors rather than by physical phenomena. The results of smoothing the variations in TEC and f_oF_2 by a 3-hour running mean is shown in Figure 4. It can be seen that the correlation is remarkably improved (correlation coefficient of 0.8). This shows that indeed the data are noisy and that the correlation between TEC and f_oF_2 is actually better than what the raw data indicate.

The seasonal dependence of the correlation coefficient between the variability in TEC and f_oF_2 is shown in Table 1. It can be seen that the correlation coefficient is better than 0.7 in winter and summer. However, during equinox, especially in April and May 1980, the correlation drops to about 0.5. The reason for the low correlation is the large post sunset enhancements in TEC which were not seen in f_oF_2 . These post sunset enhancements, observed mainly in equinox months, are attributed to electron fluxes arriving from the equatorial regions along the magnetic lines of force. They are latitude-dependent (Soicher et al., 1984) and thus did not affect the f_oF_2 observations.

The diurnal dependence of the correlation is shown in Table 2. It can be seen that higher correlation occurs during day time than night time. Between 10-18 hours local time, the correlation coefficient is about 0.7; while between 22-06 hours local time, the correlation drops to between 0.55-0.6.

TABLE 1

SEASON	MONTH	CORRELATION COEFFICIENT
Winter	Jan 80	0.74
	Feb 80	0.73
Equinox	Mar 80	0.66
	Apr 80	0.50
	May 80	0.47
Summer	Jul 81	0.73
	Jul 92	0.77
	Aug 92	0.70

TABLE 2

HOURS LT	CORRELATION COEFFICIENT
02-06	0.55
06-10	0.63
10-14	0.70
14-18	0.71
18-22	0.60
22-02	0.61

The ionosphere is known to vary substantially with magnetic activity (Goodman, 1992). To ascertain whether magnetic activity has any impact on the cross correlation of f_oF_2 and TEC, two time periods - one quiet (18-26 January 1980) and one active (14-19 February 1980) - were examined. The former with index $Ap < 6$ is depicted in Figure 5 with maximum correlation coefficient of 0.823 and the latter with index $Ap < 40$ is depicted in Figure 6 with maximum correlation coefficient 0.774. Thus it appears that magnetic changes do not have a marked impact on the correlation of the two ionospheric parameters.

The slab thickness τ is determined from the ratio $TEC/(f_oF_2)^2$. It is expected that the variability of τ will be smaller than the variability of either the TEC or f_oF_2 because of the good correlation between these two ionospheric parameters. Thus, global models of the slab thickness updated with real-time measurements of TEC obtained with the GPS network, might give improved values

of f_oF_2 . However, the calculated variability in τ is of the same order as the corresponding variability of TEC or f_oF_2 . For example, for the period 11-29 February 1980, the standard deviation of the hourly variability of τ is 16% in comparison to that of TEC which is 22%.

3. CONCLUSIONS

The high cross-correlation (>0.75) coefficient for f_oF_2 and TEC in the limited data presented in this paper raises the possibility that real time TEC measurements may be used to update f_oF_2 value determinations. The cross-correlation may even be higher if the geographic subionospheric point of the TEC measurement is closer to the geographic point of the f_oF_2 measurement, which introduces an error, in addition to the possible inherent measurement uncertainties. TEC measurements utilizing satellite emitted signals are passive in nature and do not burden the electromagnetic spectrum. In addition, the availability of the global GPS constellation to provide instantaneous time-delay, or equivalently TEC, values could provide an instantaneous updating of f_oF_2 models on a global basis as well as on a regional basis. Such capability is important for HF communication along short, medium and long range paths.

4. REFERENCES

- Goodman, J. M., HF Communication Science and Technology, Van Nostrand Reinhold, New York, NY, PP631, 1992.
- Klobuchar, J. A., Ionospheric Time-Delay Algorithm for Single Frequency GPS Users, IEEE Trans AES, AES-23, 325-331, 1987.
- Soicher, H., Z. Houminer and A. Shuval, Total Electron content in the Middle East, Radio Science, 17, 1623-1631, 1982.
- Soicher, H., J. A. Klobuchar and P. H. Doherty, Spatial Variability of Total Electron Content in the Eastern Mediterranean Region, Radio Science, 19, 757-764, 1984.
- Wilson, B. D. and A. J. Mannucci, Instrumental Biases in Ionospheric Measurements Derived from GPS Data, Proceedings of the Institute of Navigation GPS-93, Salt Lake City, Utah, September 1993.

5. FIGURE CAPTIONS

1. Hourly values of the variability of f_oF_2 (at Cyprus) and TEC (at Israel) for the time period 1-16 August 1992.
2. Crosscorrelation function for the f_oF_2 and TEC values of Figure 1.
3. Autocorrelation functions for the f_oF_2 and TEC values of Figure 1.
4. Smoothed hourly values of the variability values of Figure 1.
5. The variability (in percent) of f_oF_2 and TEC for the geomagnetically quiet period of 18-26 January 1980.
6. The variability (in percent) of f_oF_2 and TEC for the geomagnetically active period of 14-19 February 1980.

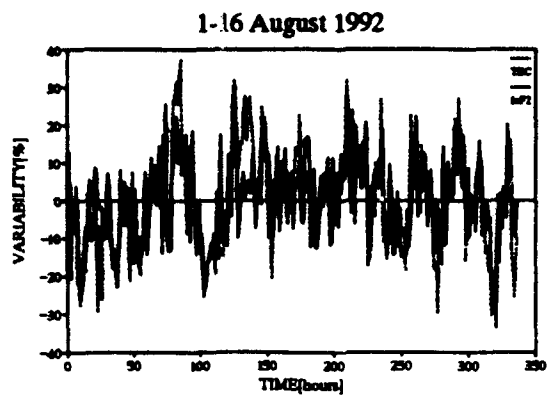


Figure 1

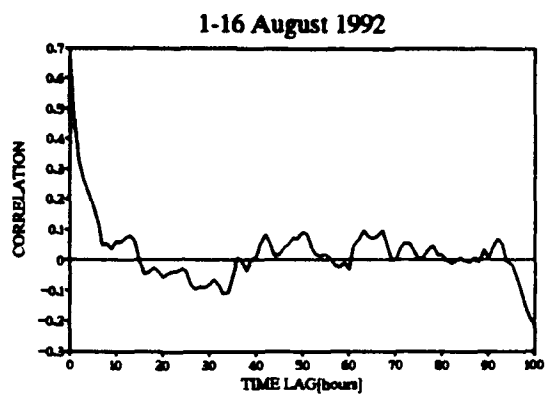


Figure 2

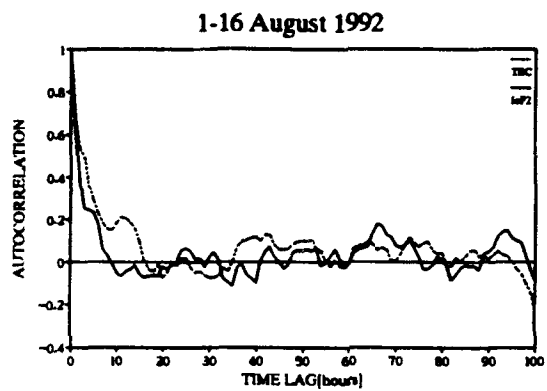


Figure 3

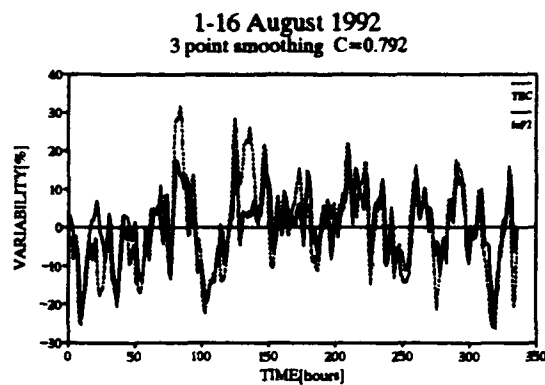


Figure 4

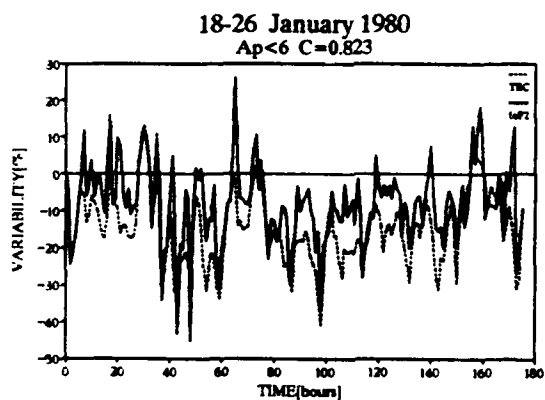


Figure 5

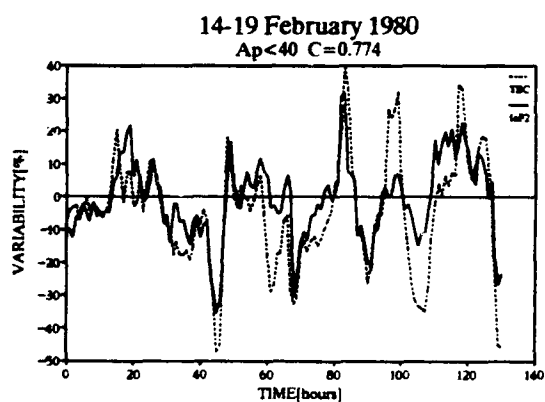


Figure 6

APPENDIX B

Experimental results for the period January - May 1980,
July 1981 and June - August 1992.

The following are the figure captions:

- Fig B-1 Hourly values of the variability of foF2 and TEC
for the period 8-31 January 1980.
- Fig B-2 Cross-correlation function between foF2 and TEC
for the period 8-31 January 1980.
- Fig B-3 Auto-correlation functions for foF2 and TEC for
the period 8-31 January 1980.
- Fig B-4 Hourly values of the variability of τ for the
period 8-31 January 1980.
- Fig B-5 Hourly values of the variability of foF2 and TEC
for the period 11-29 February 1980.
- Fig B-6 Cross-correlation function between foF2 and TEC
for the period 11-29 February 1980.
- Fig B-7 Auto-correlation functions for foF2 and TEC for
the period 11-29 February 1980.
- Fig B-8 Hourly values of the variability of τ for the
period 11-29 February 1980.
- Fig B-9 Hourly values of the variability of foF2 and TEC
for the period 1-31 March 1980.
- Fig B-10 Cross-correlation function between foF2 and TEC
for the period 1-31 March 1980.
- Fig B-11 Auto-correlation functions for foF2 and TEC for
the period 1-31 March 1980.
- Fig B-12 Hourly values of the variability of τ for the =
period 1-31 March 1980.
- Fig B-13 Hourly values of the variability of foF2 and TEC
for the period 1-30 April 1980.
- Fig B-14 Cross-correlation function between foF2 and TEC
for the period 1-30 April 1980.
- Fig B-15 Auto-correlation functions for foF2 and TEC for
the period 1-30 April 1980.

- Fig B-16 Hourly values of the variability of r for the period 1-30 April 1980.
- Fig B-17 Hourly values of the variability of foF2 and TEC for the period 1-31 May 1980.
- Fig B-18 Cross-correlation function between foF2 and TEC for the period 1-31 May 1980.
- Fig B-19 Auto-correlation functions for foF2 and TEC for the period 1-31 May 1980.
- Fig B-20 Hourly values of the variability of r for the period 1-31 May 1980.
- Fig B-21 Hourly values of the variability of foF2 and TEC for the period 2-31 July 1981.
- Fig B-22 Cross-correlation function between foF2 and TEC for the period 2-31 July 1981.
- Fig B-23 Auto-correlation functions for foF2 and TEC for the period 2-31 July 1981.
- Fig B-24 Hourly values of the variability of r for the period 2-31 July 1981.
- Fig B-25 Hourly values of the variability of foF2 and TEC for the period 1-30 June 1992.
- Fig B-26 Smoothed hourly value by a 3 hour running mean for the period 1-30 June 1992.
- Fig B-27 Smoothed hourly value by a 5 hour running mean for the period 1-30 June 1992.
- Fig B-28 Cross-correlation function between foF2 and TEC for the period 1-30 June 1992.
- Fig B-29 Auto-correlation functions for foF2 and TEC for the period 1-30 June 1992.
- Fig B-30 Hourly values of the variability of r for the period 1-30 June 1992.
- Fig B-31 Hourly values of the variability of foF2 and TEC for the period 2-31 July 1992.
- Fig B-32 Smoothed hourly value by a 3 hour running mean for the period 2-31 July 1992.
- Fig B-33 Smoothed hourly value by a 5 hour running mean for the period 2-31 July 1992.

- Fig B-34** **Cross-correlation function between foF2 and TEC for the period 2-31 July 1992.**
- Fig B-35** **Auto-correlation functions for foF2 and TEC for the period 2-31 July 1992.**
- Fig B-36** **Hourly values of the variability of r for the period 2-31 July 1992.**
- Fig B-37** **Hourly values of the variability of foF2 and TEC for the period 1-16 August 1992.**
- Fig B-38** **Smoothed hourly value by a 3 hour running mean for the period 1-16 August 1992.**
- Fig B-39** **Smoothed hourly value by a 5 hour running mean for the period 1-16 August 1992.**
- Fig B-40** **Cross-correlation function between foF2 and TEC for the period 1-16 August 1992.**
- Fig B-41** **Auto-correlation functions for foF2 and TEC for the period 1-16 August 1992.**
- Fig B-42** **Hourly values of the variability of r for the period 1-16 August 1992.**

8-31 January 1980
 $C=0.738$

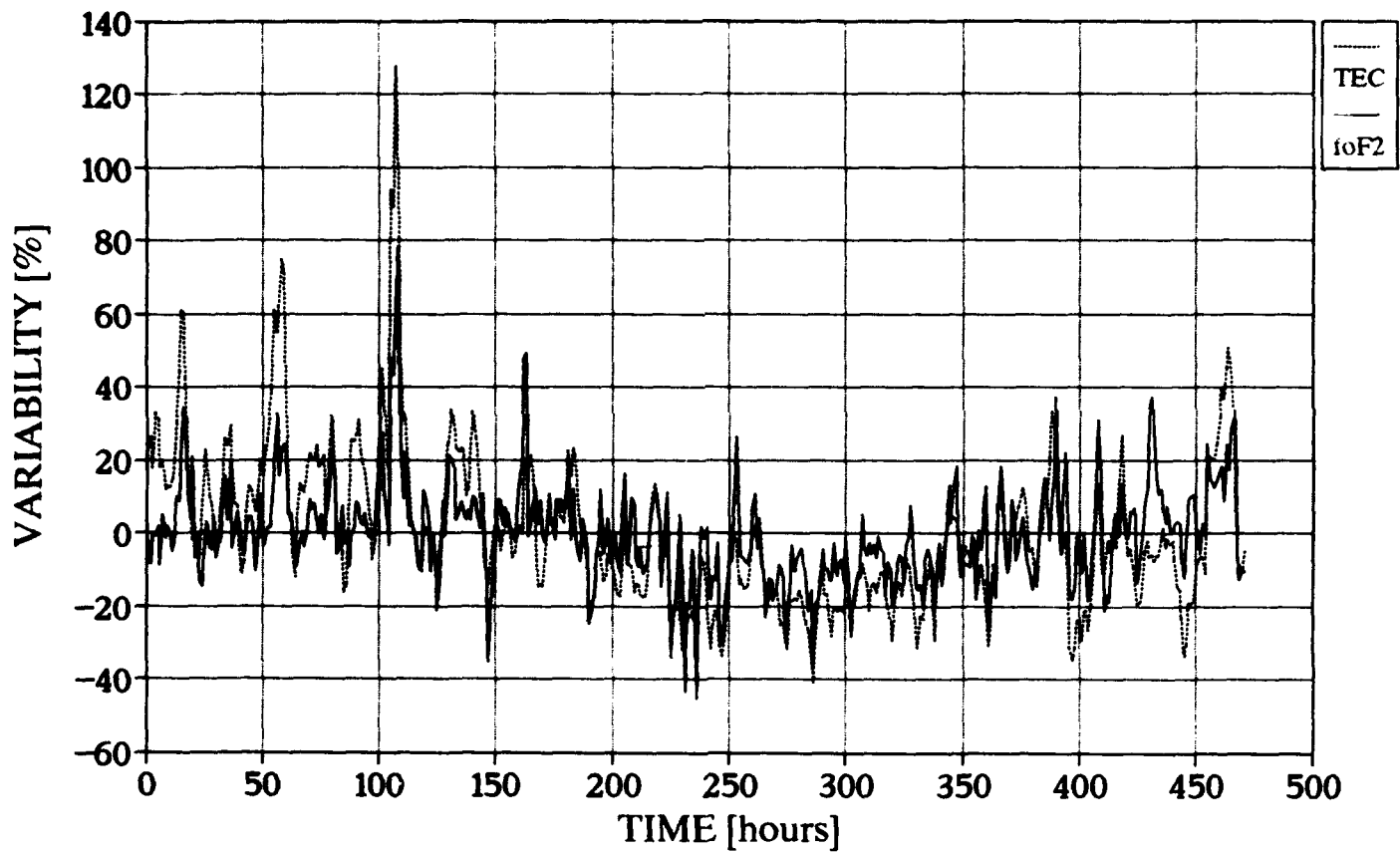


Fig B-1

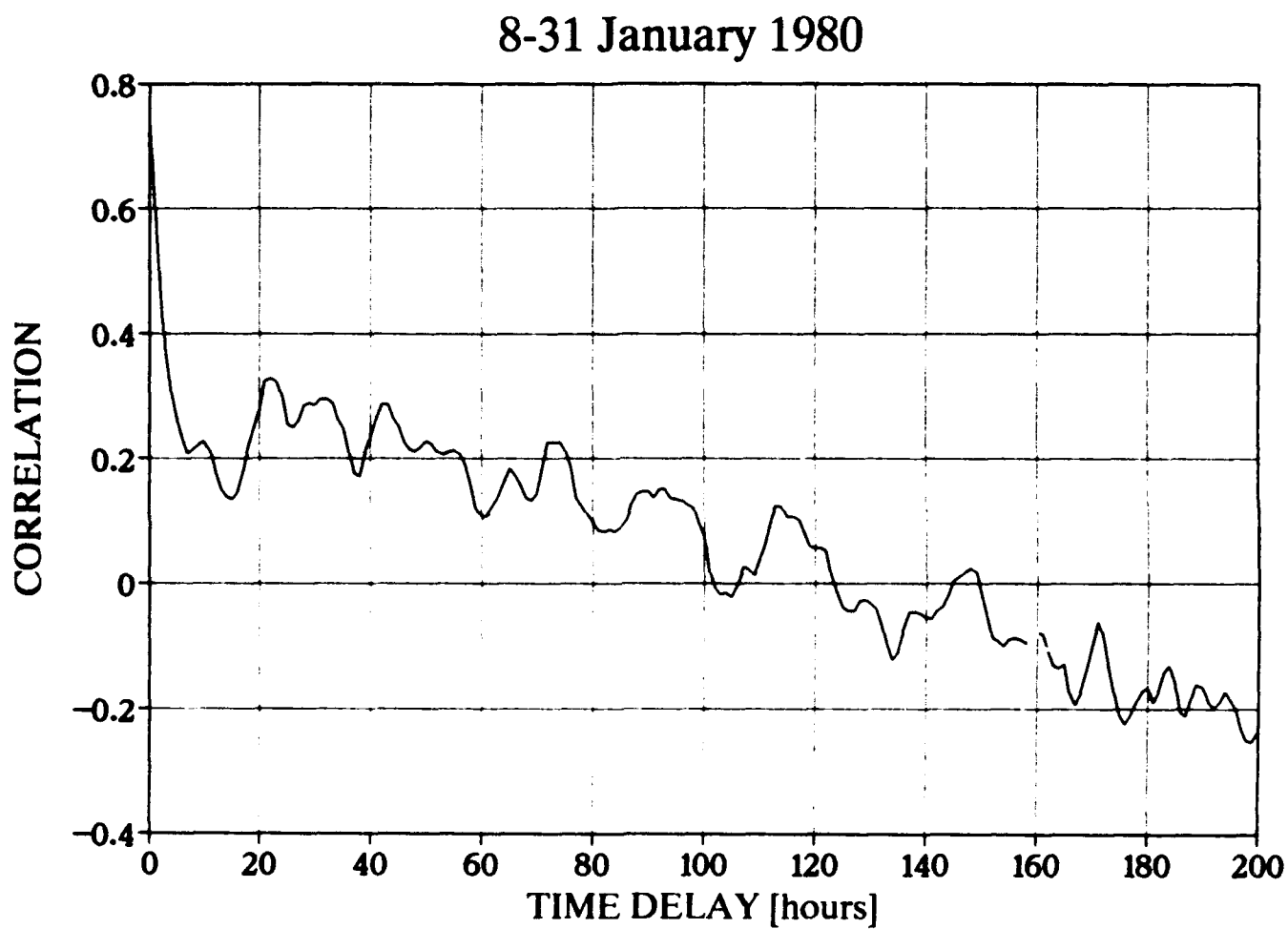


Fig B-2

8-31 January 1980

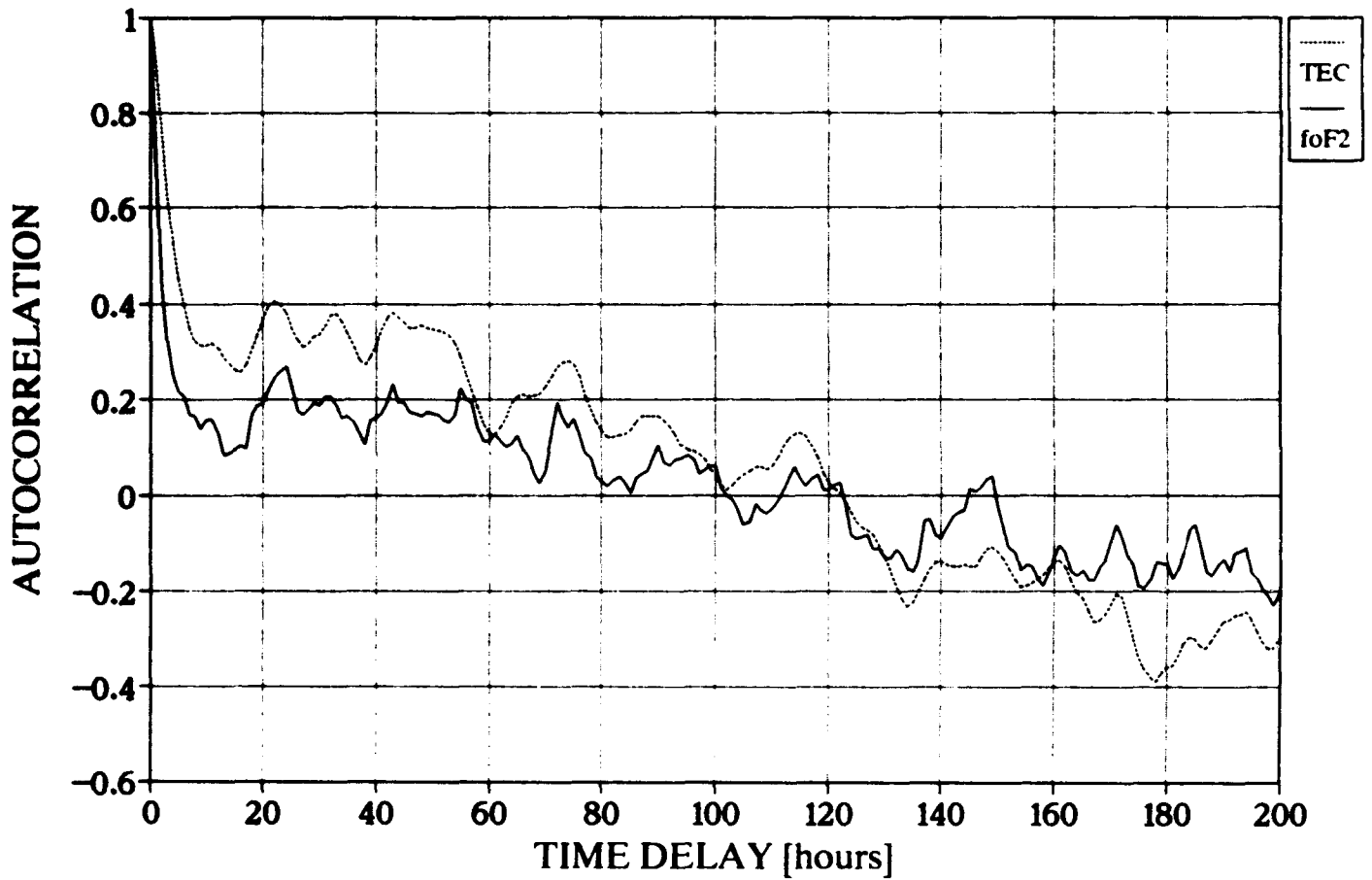


Fig B-5

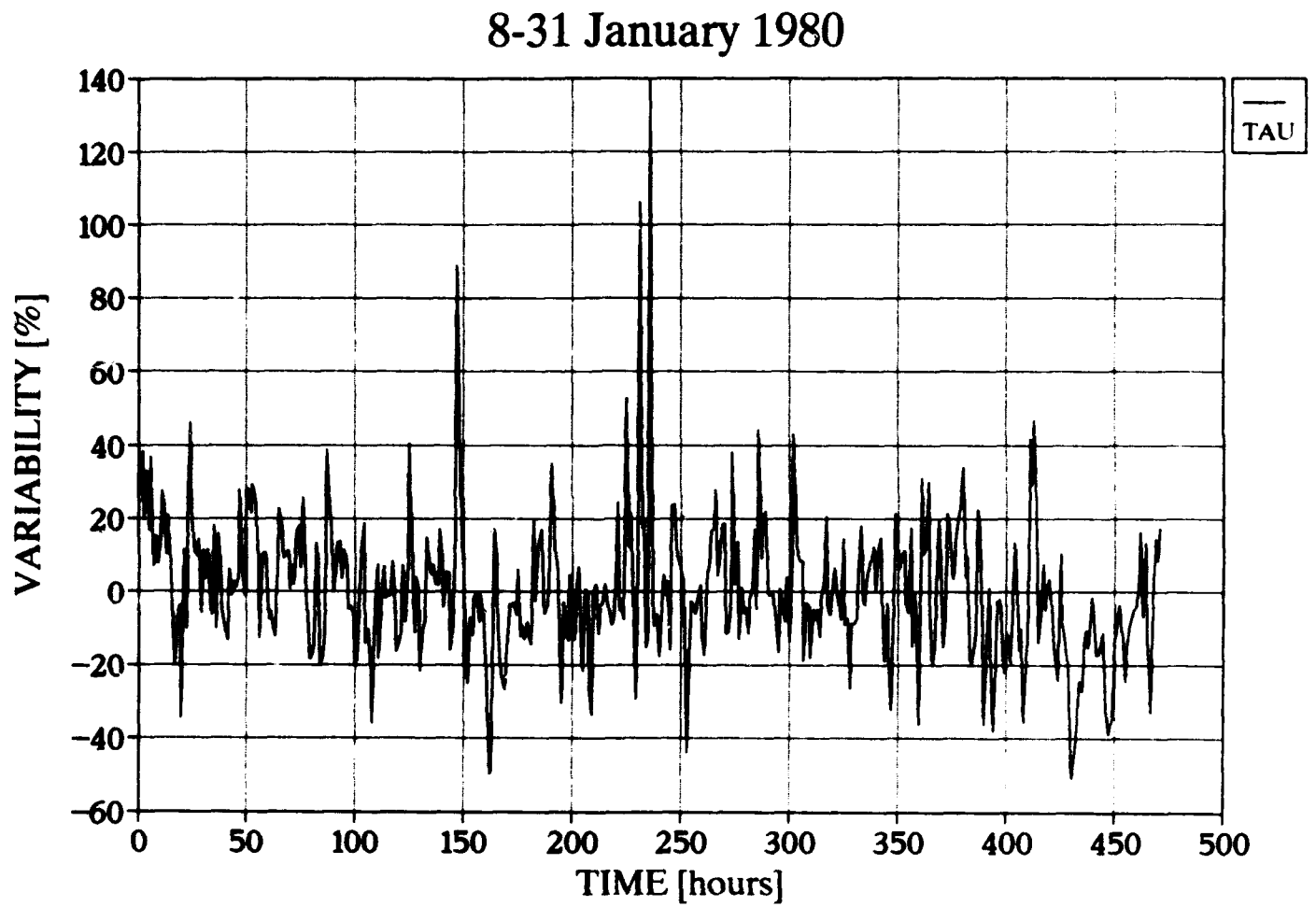


Fig B-4

11-29 February 1980
 $C = 0.733$

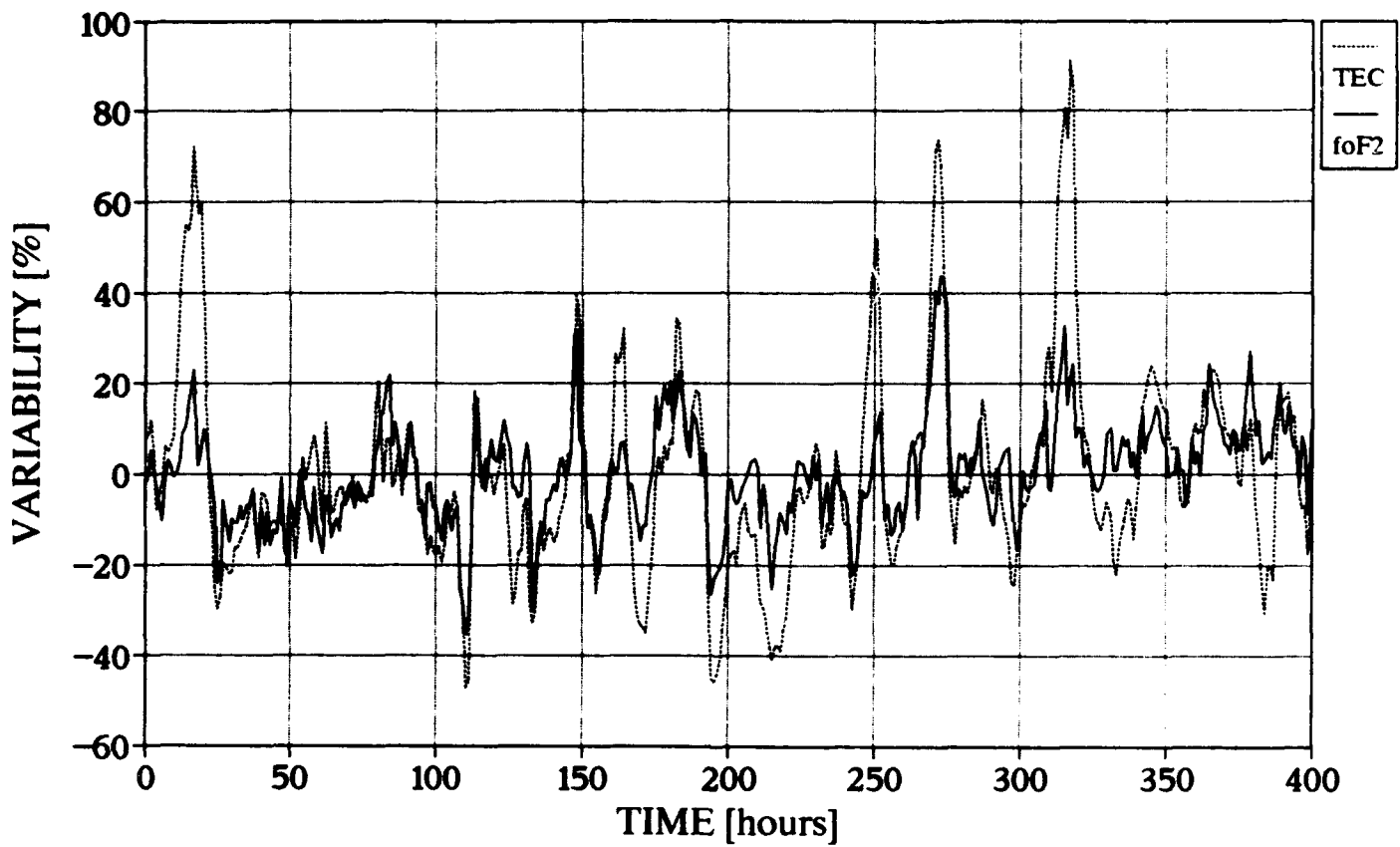


Fig B-5

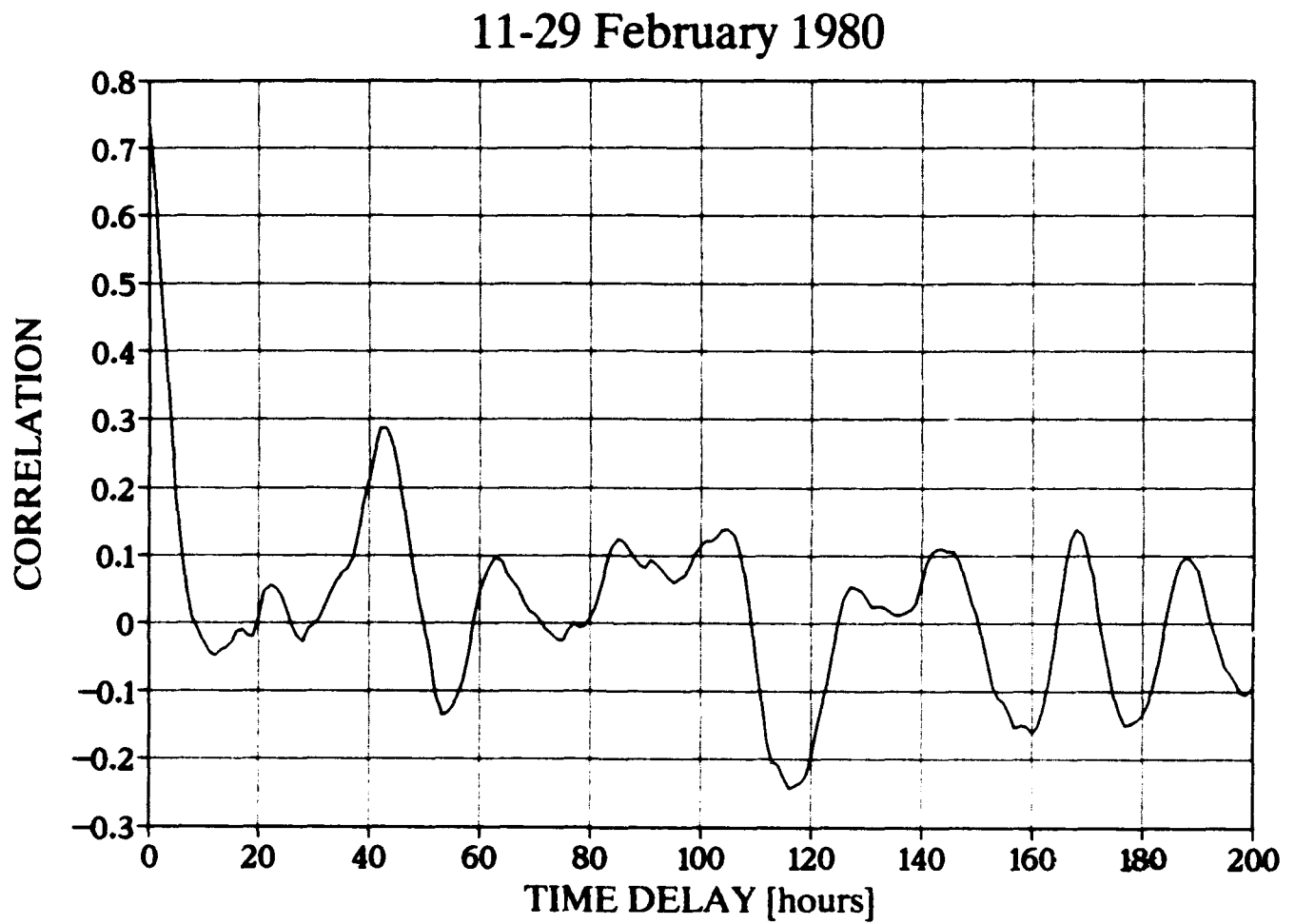


Fig B-6

11-29 February 1980

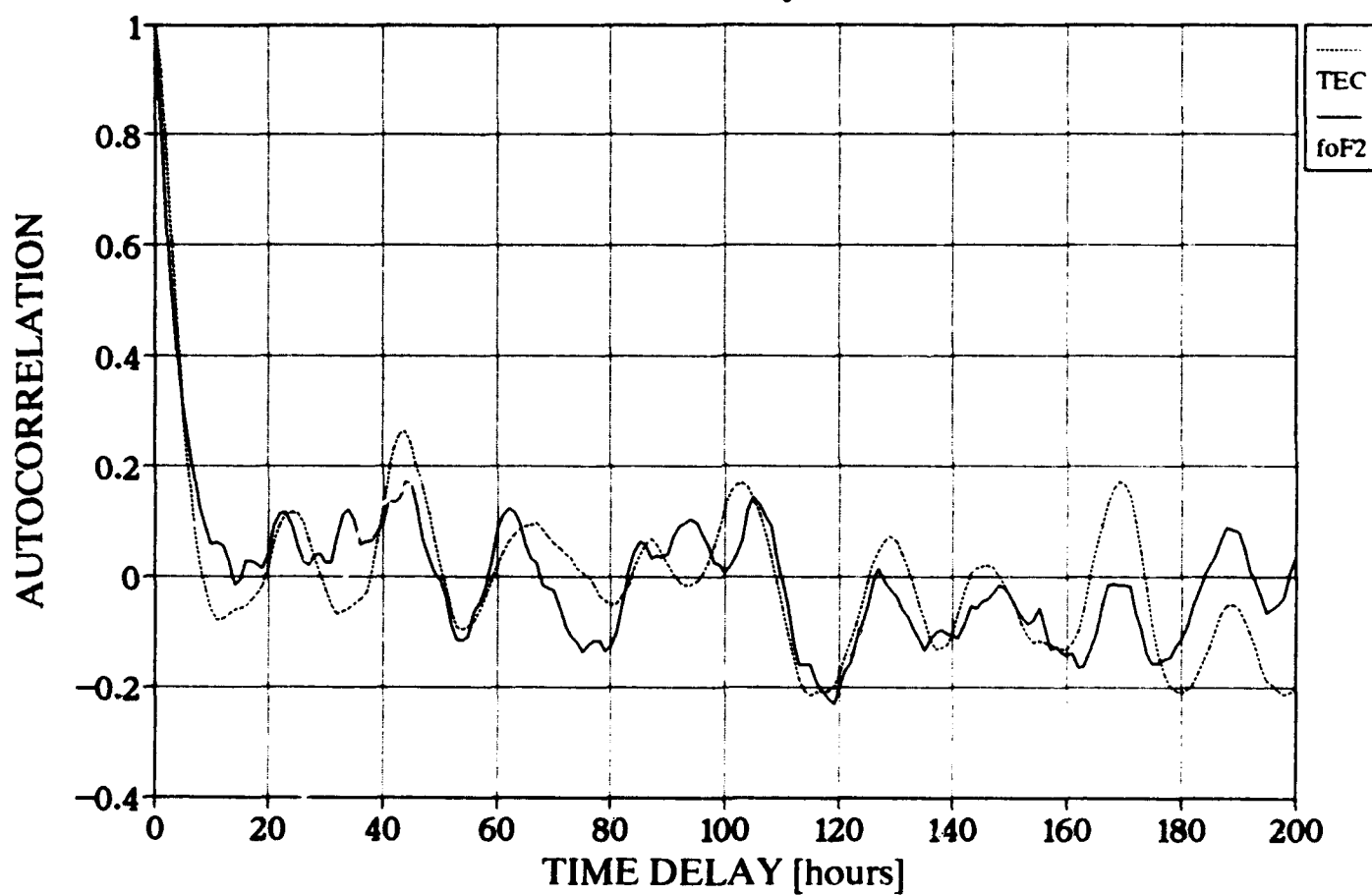


Fig B-7

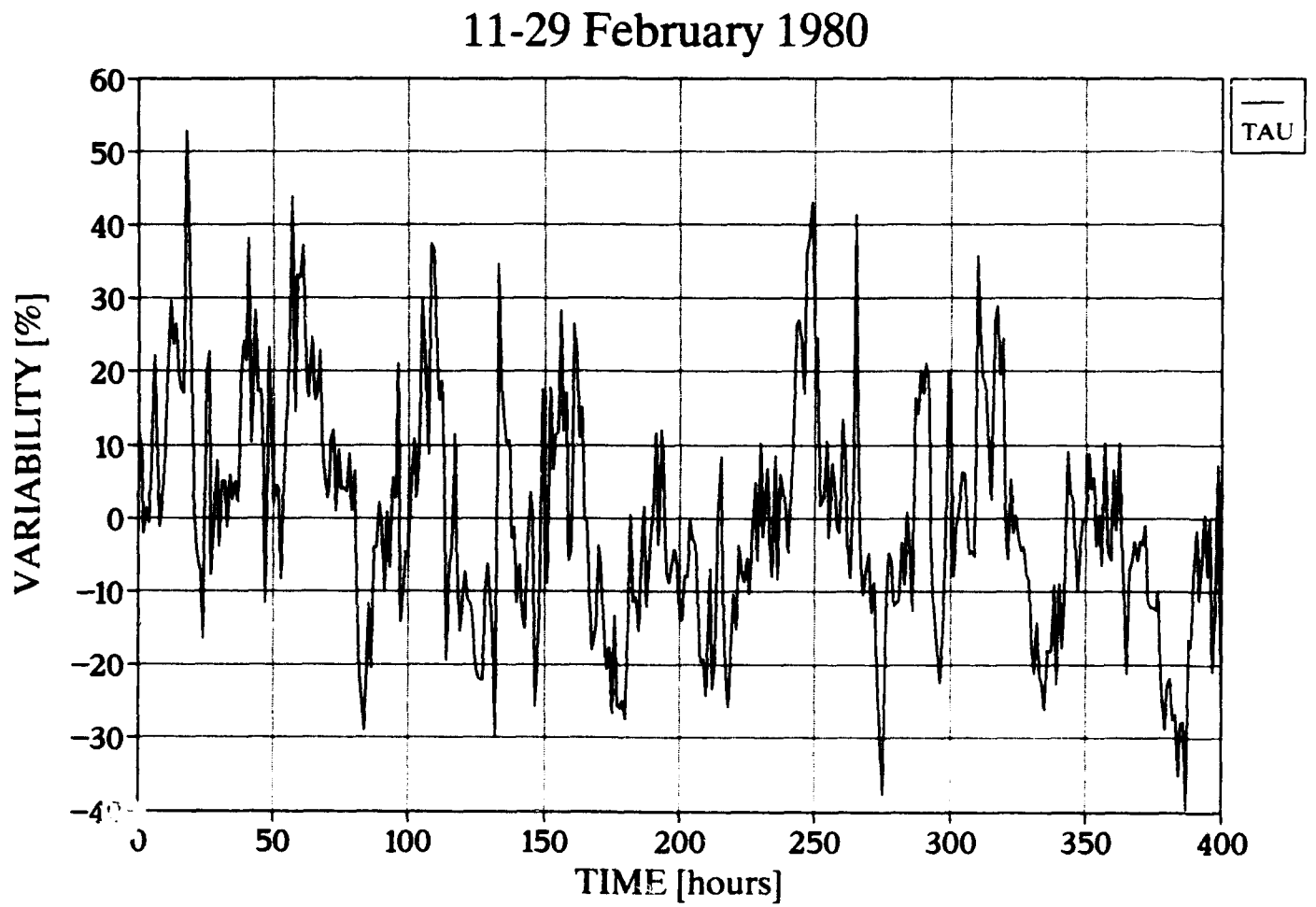


Fig B-8

1-31 march 1980

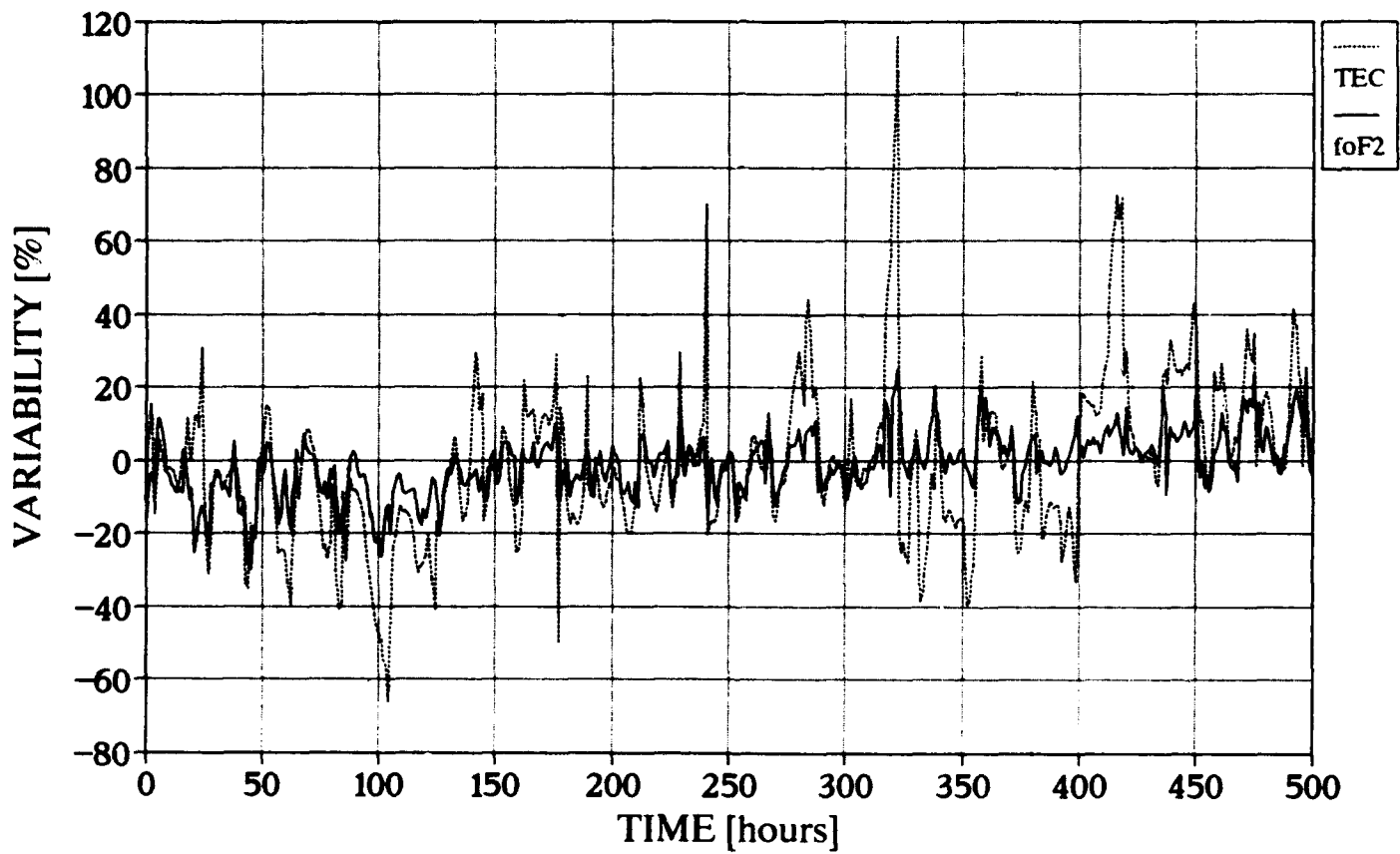
 $C=0.661$ 

Fig B-9

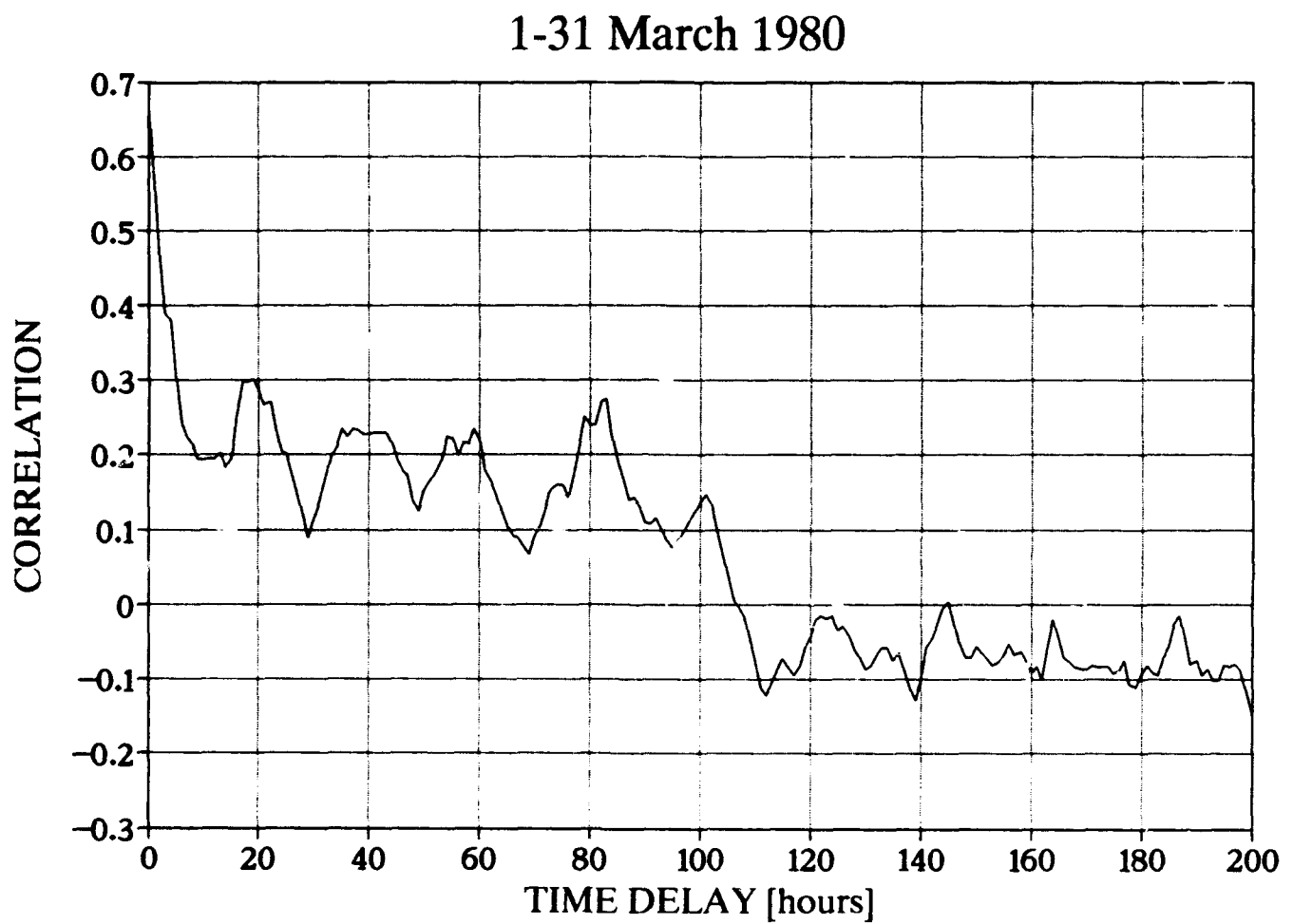


Fig B-10

1-31 March 1980

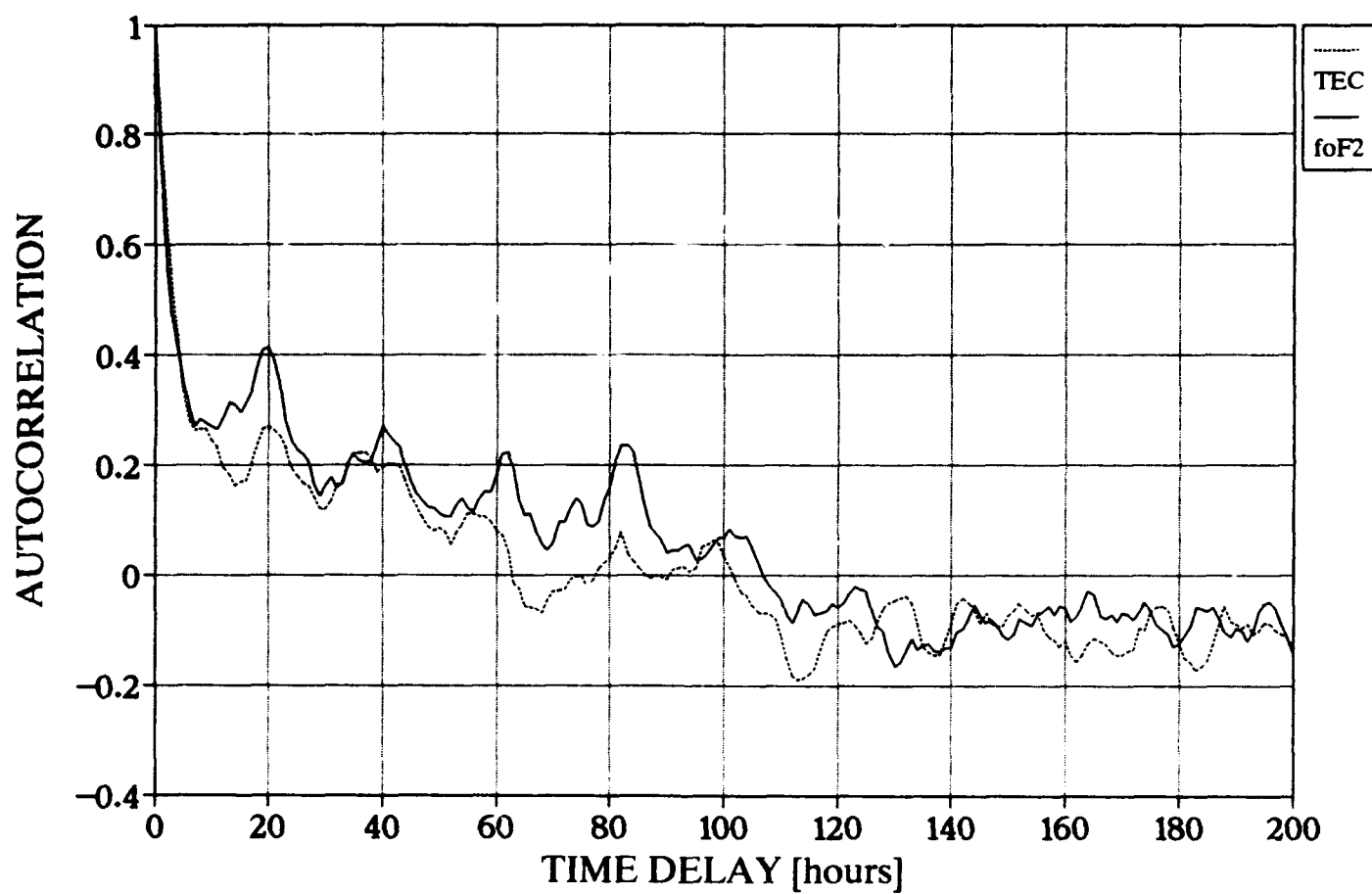


Fig B-11

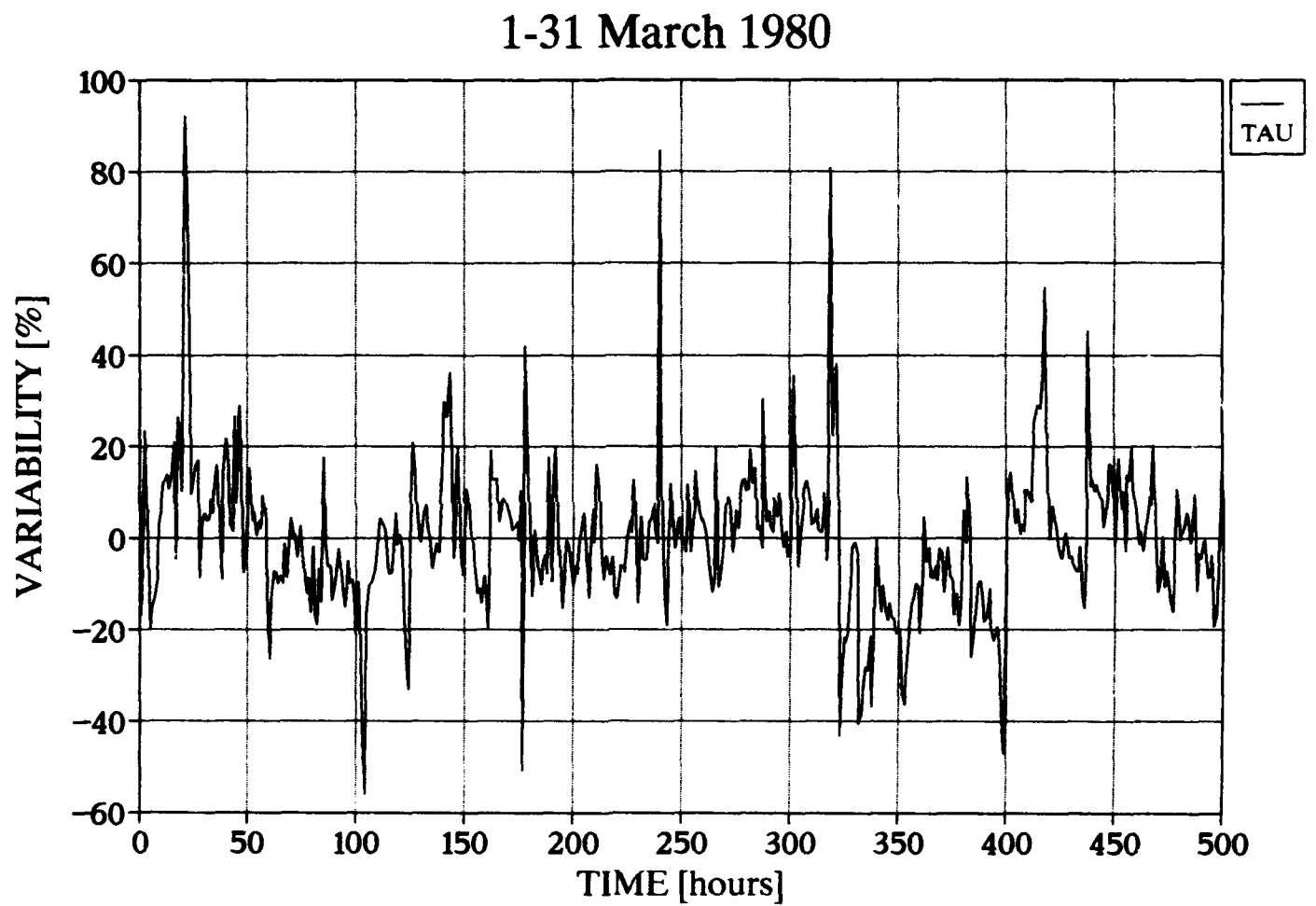


Fig B-12

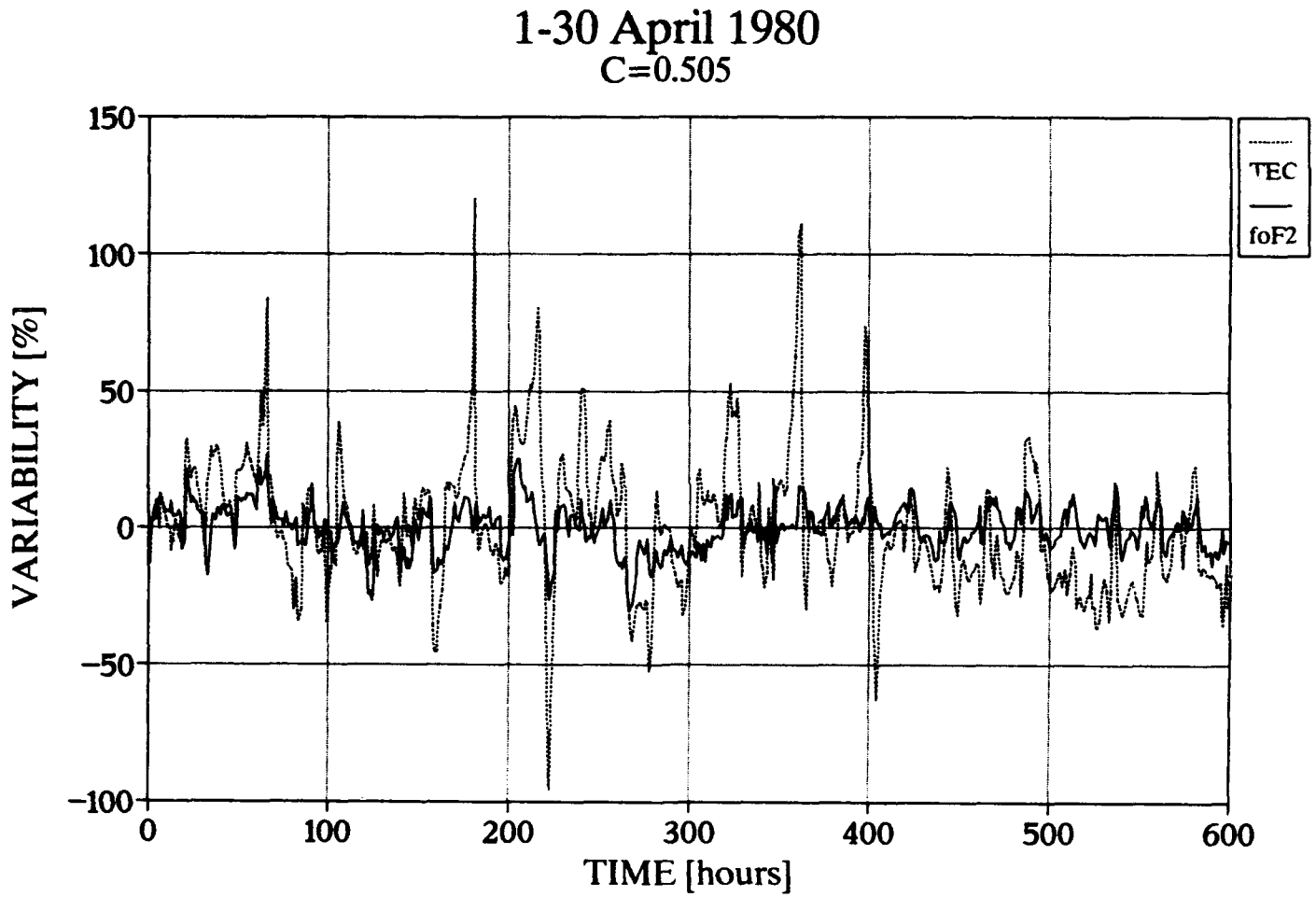


Fig B-15

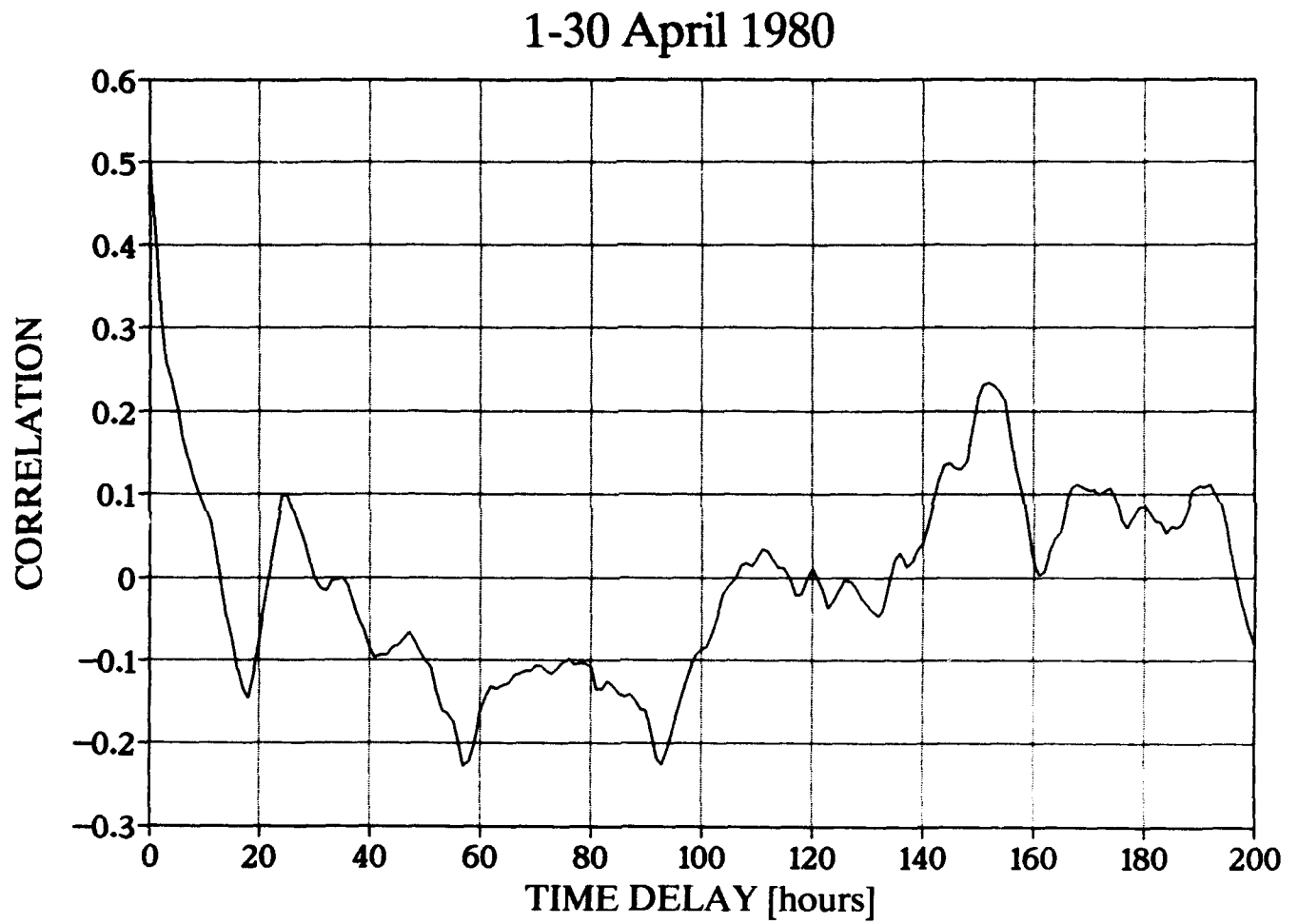


Fig B-14

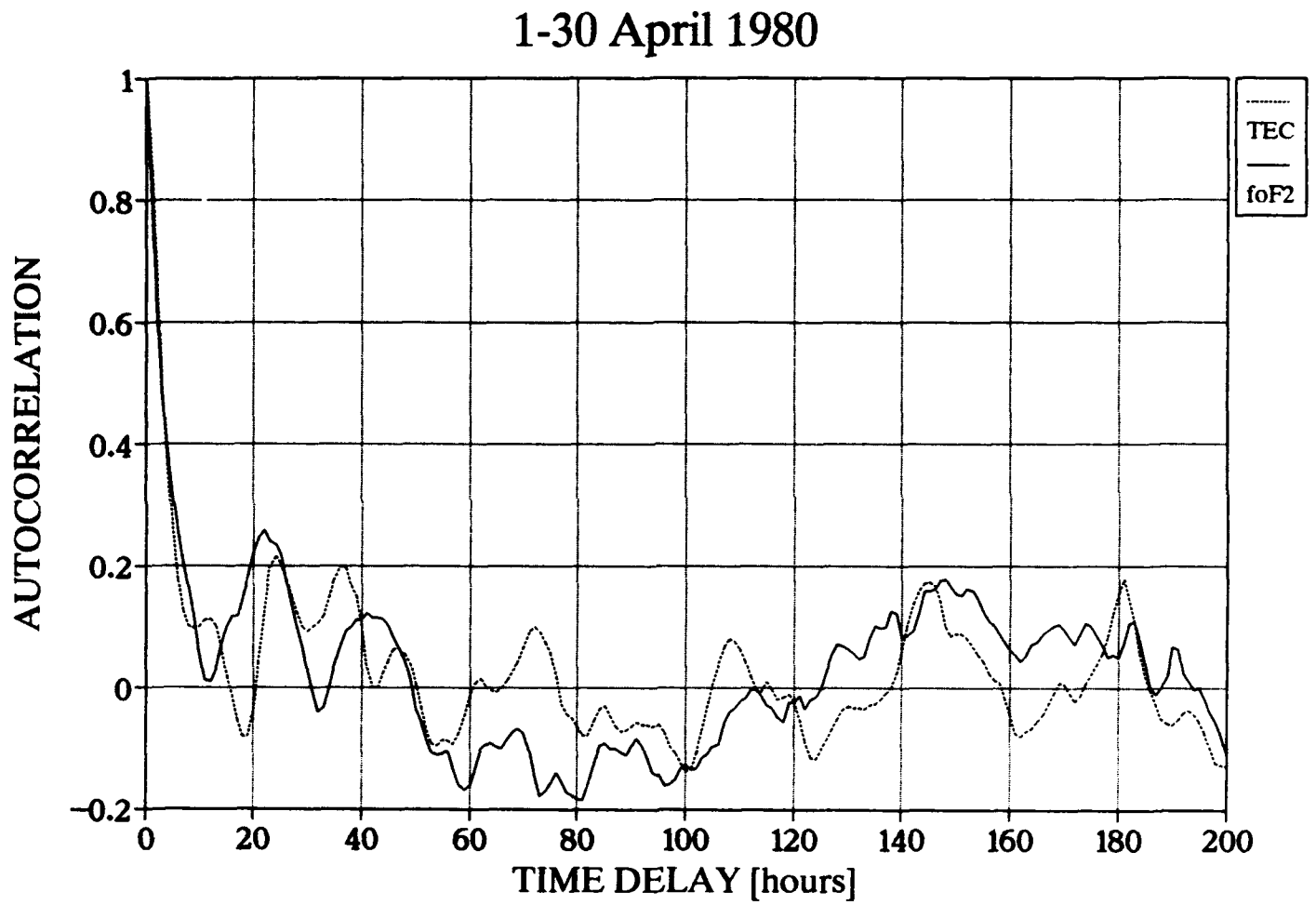


Fig B-15

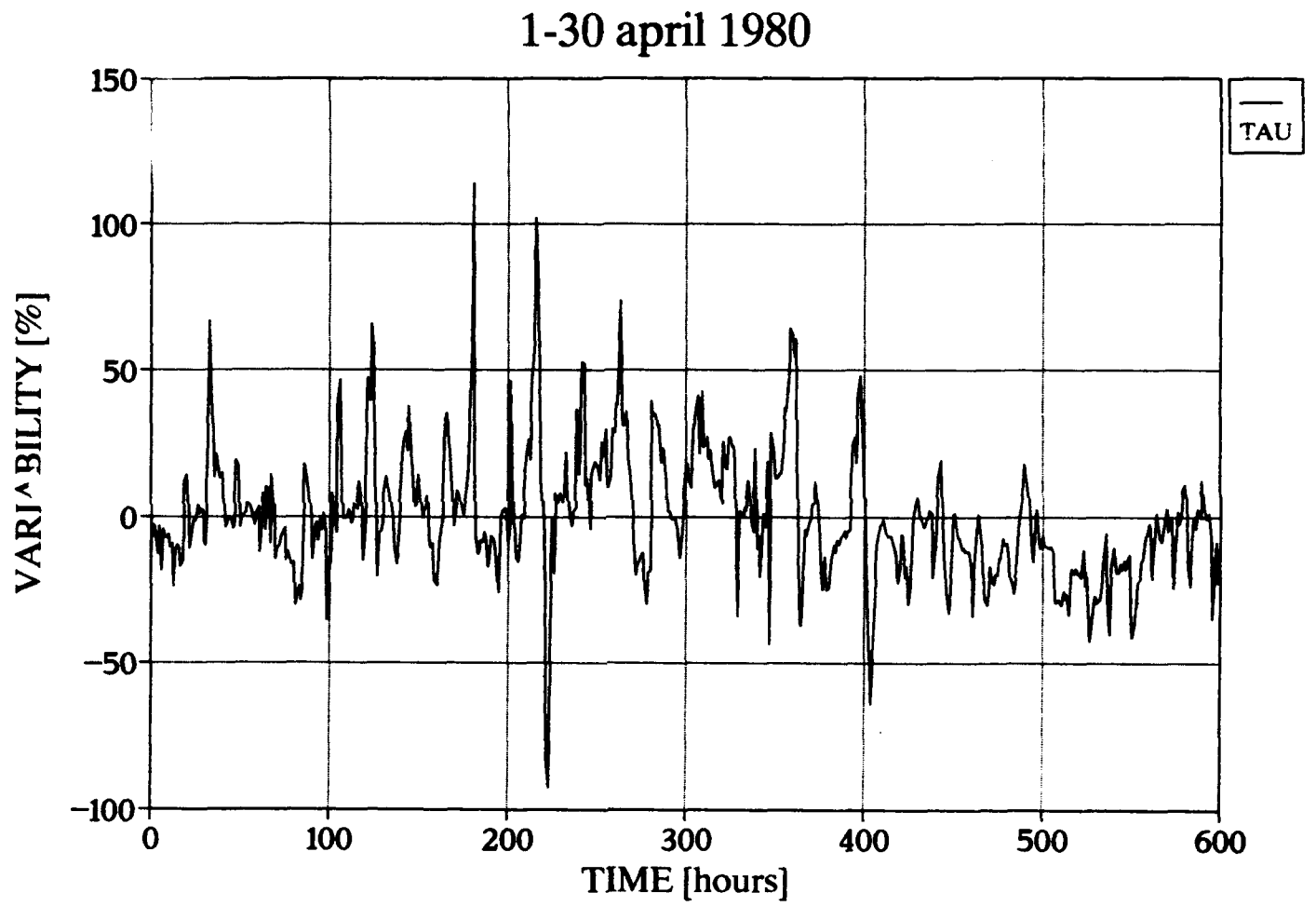


Fig B-16

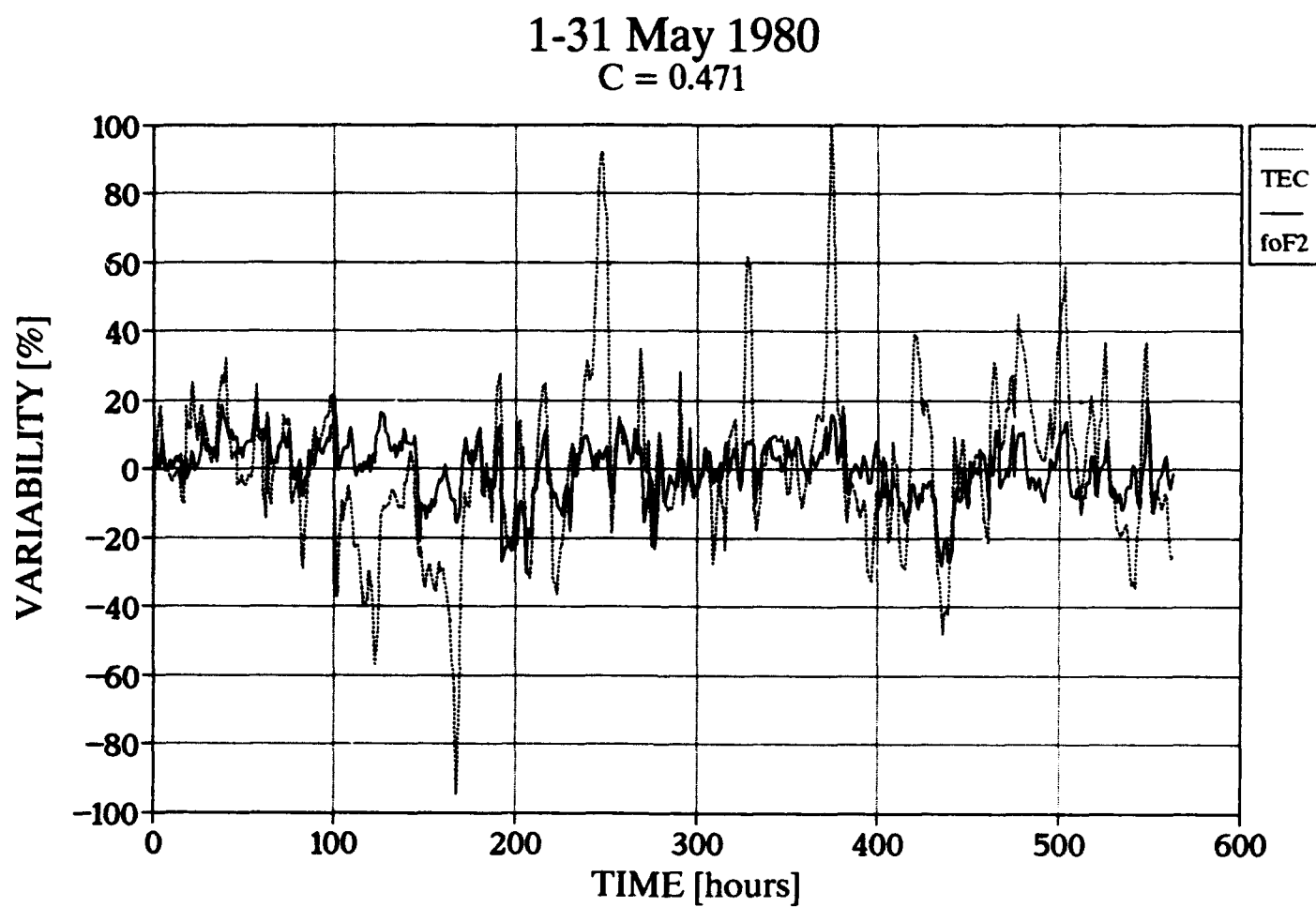


Fig B-17

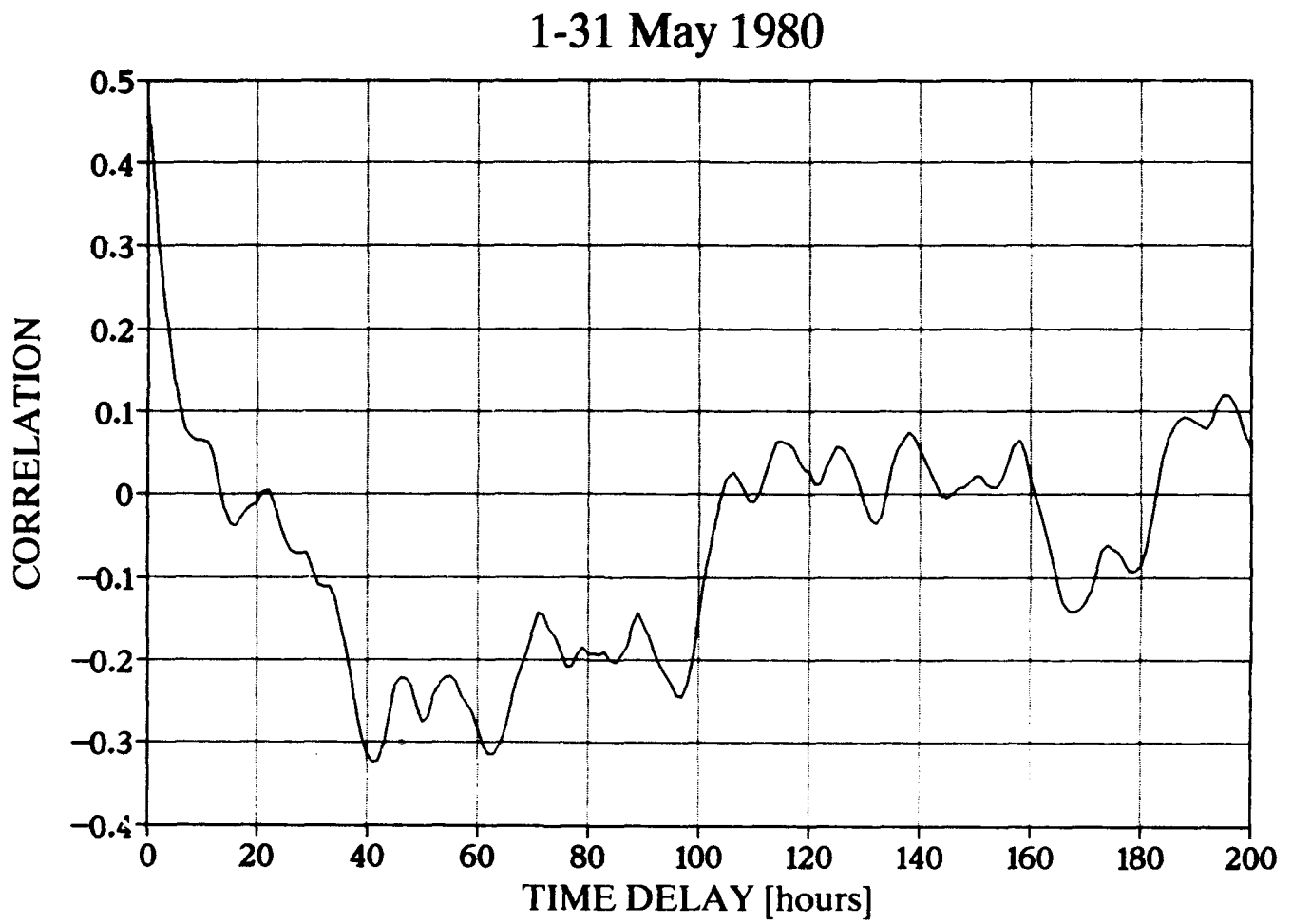


Fig B-18

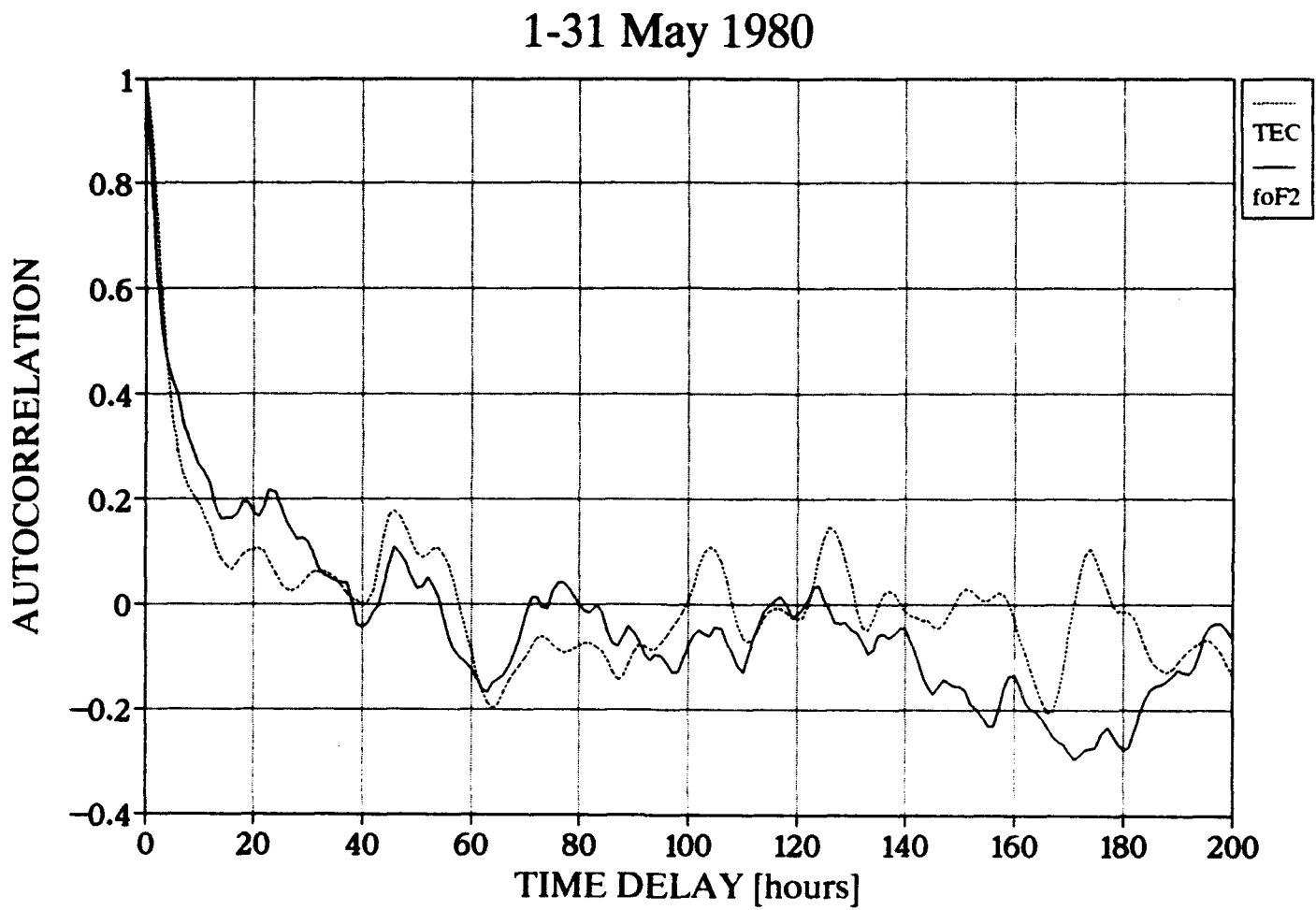


Fig B-19

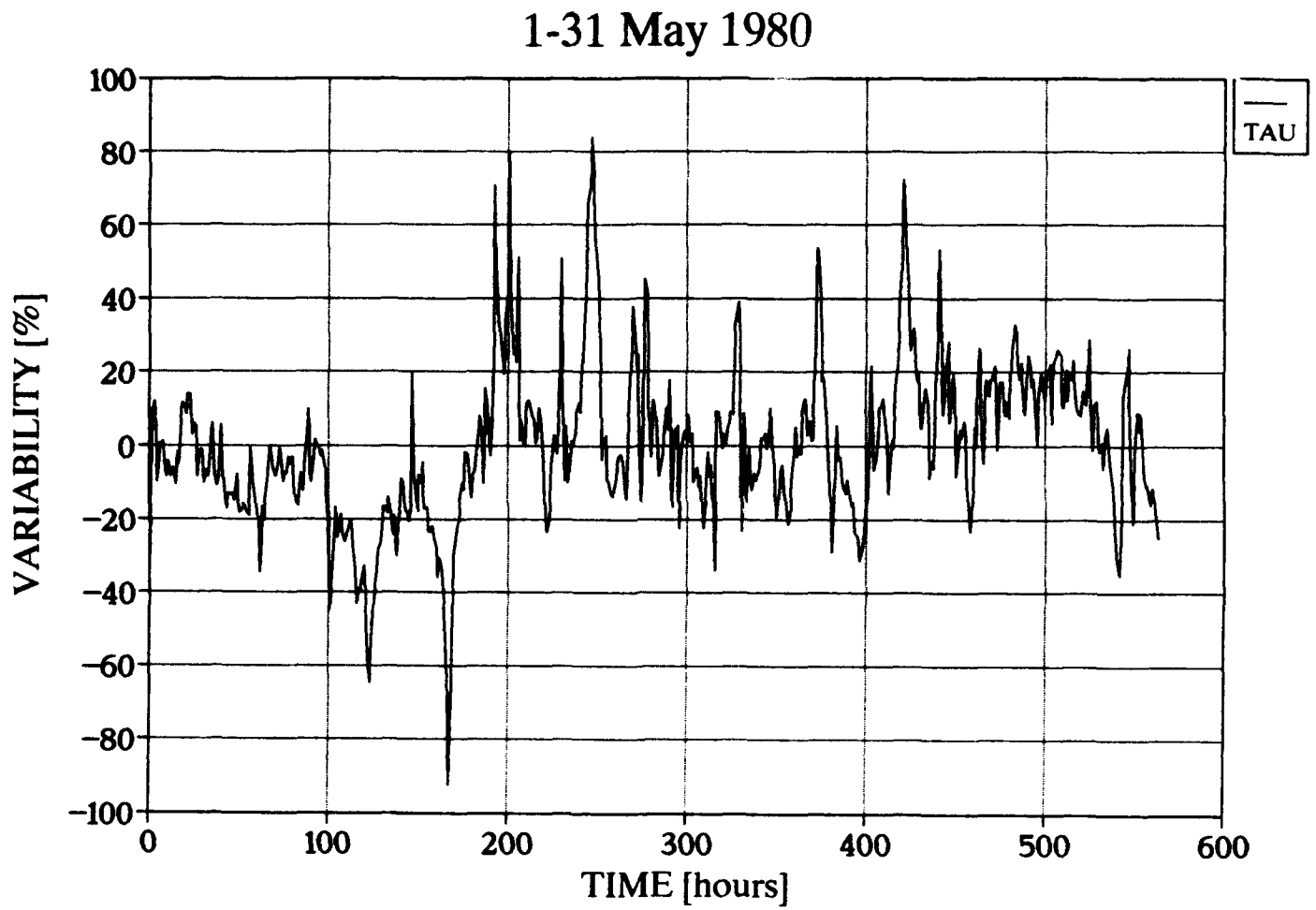


Fig B-20

2-31 July 1981
C=0.730

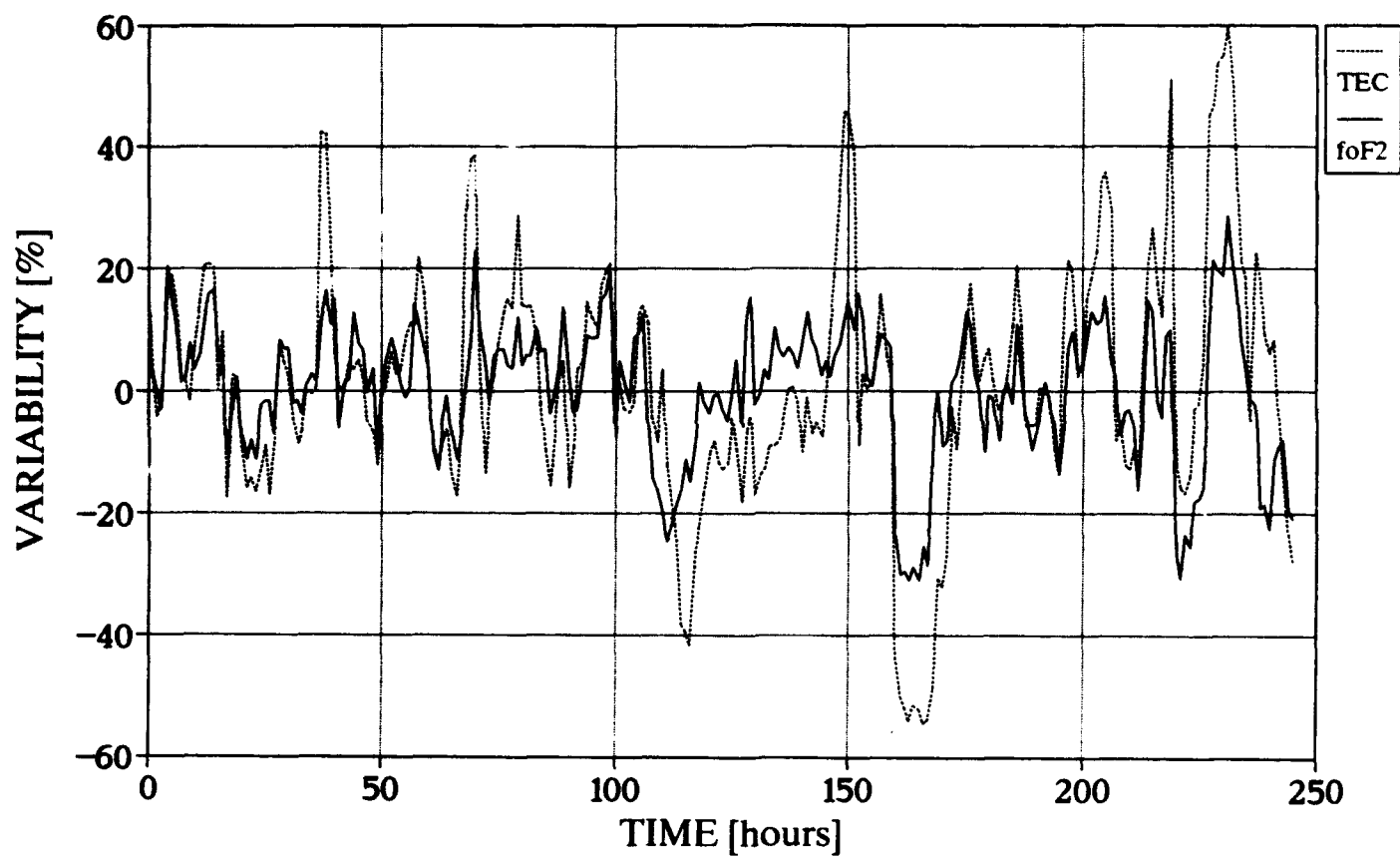


Fig B-21

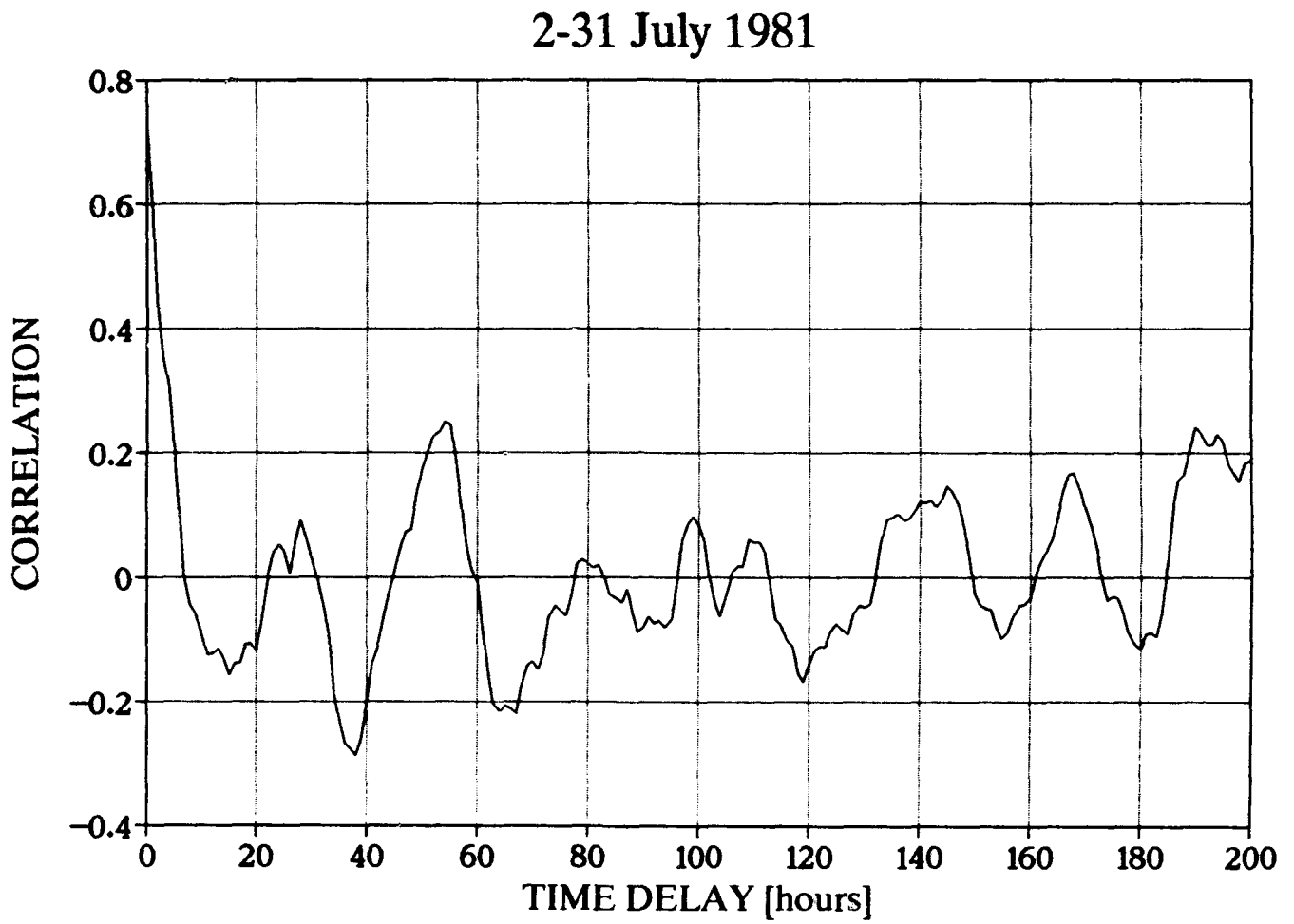


Fig B-22

2-31 July 1981

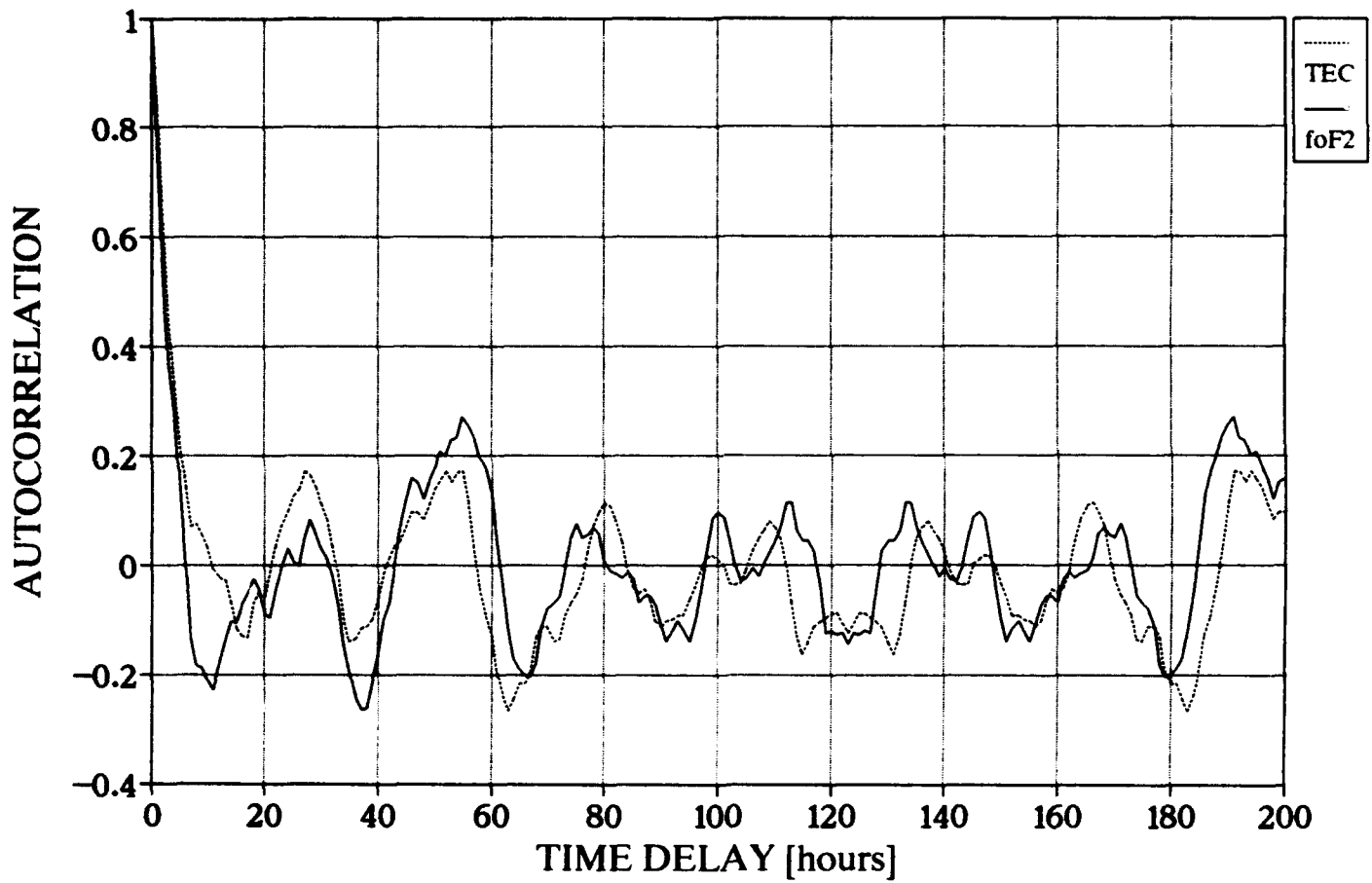


Fig B-25

2-31 July 1981

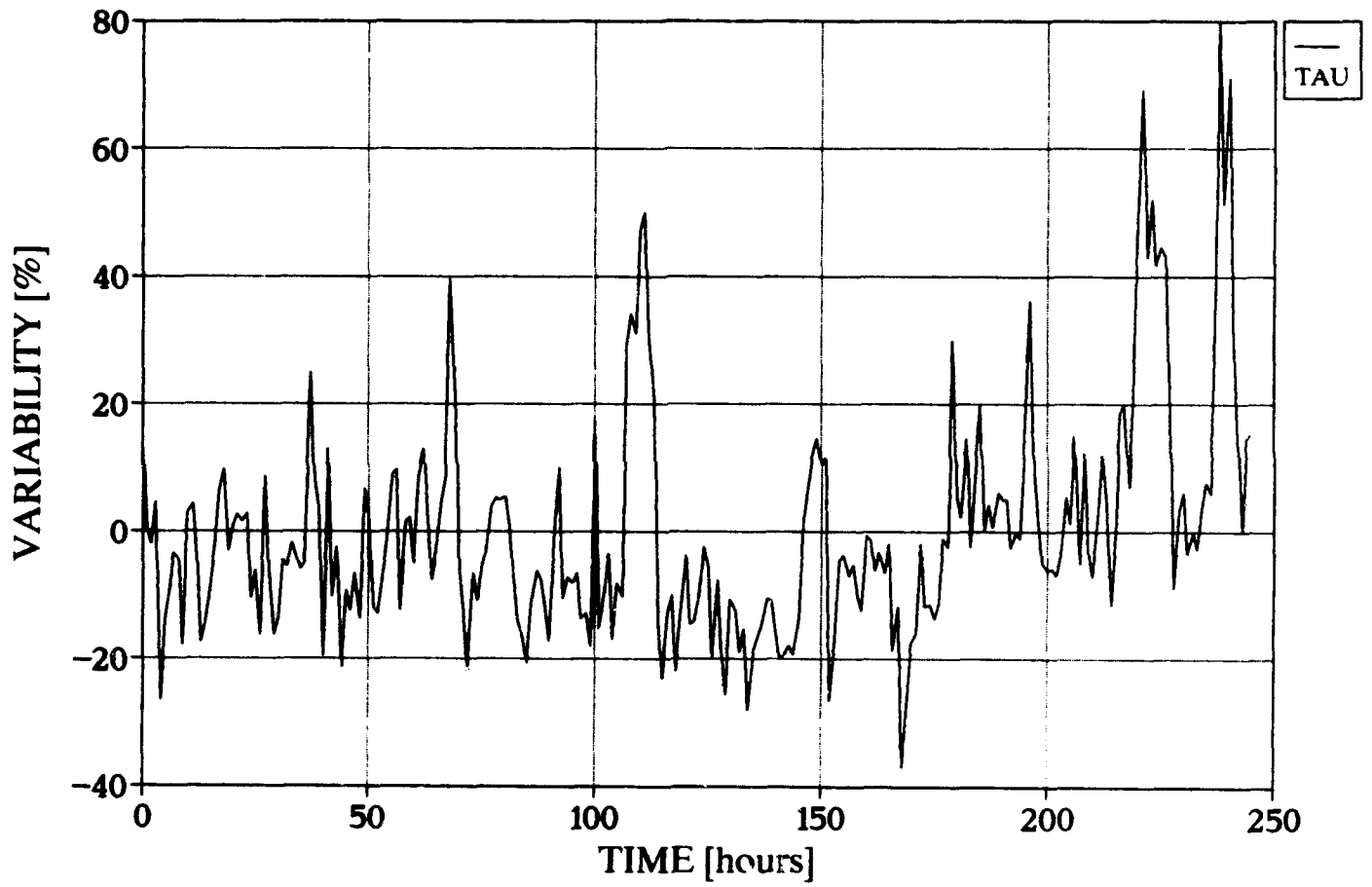


Fig B-24

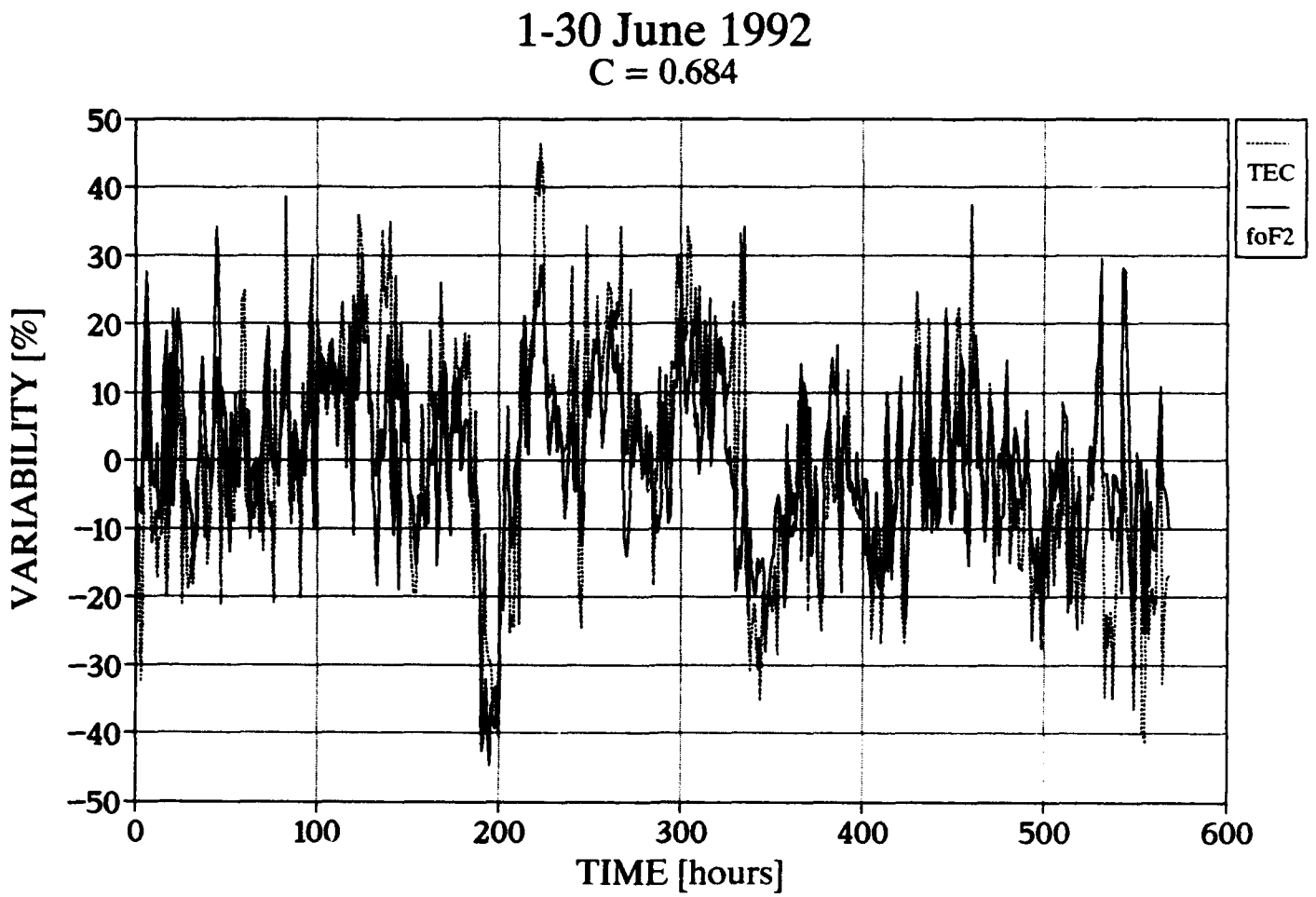


Fig B-25

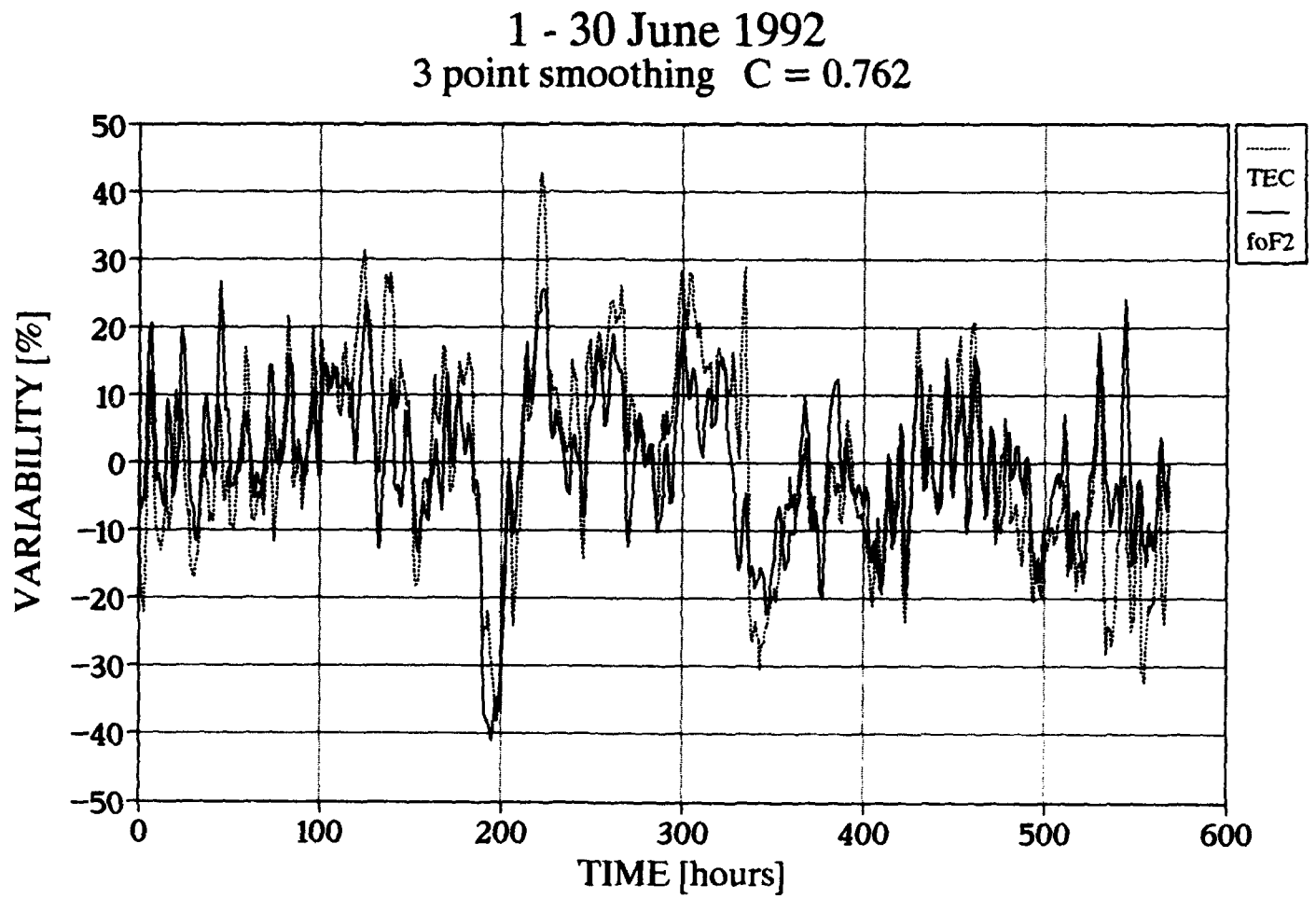


Fig B-26

1 - 30 June 1992
5 point smoothing $C = 0.784$

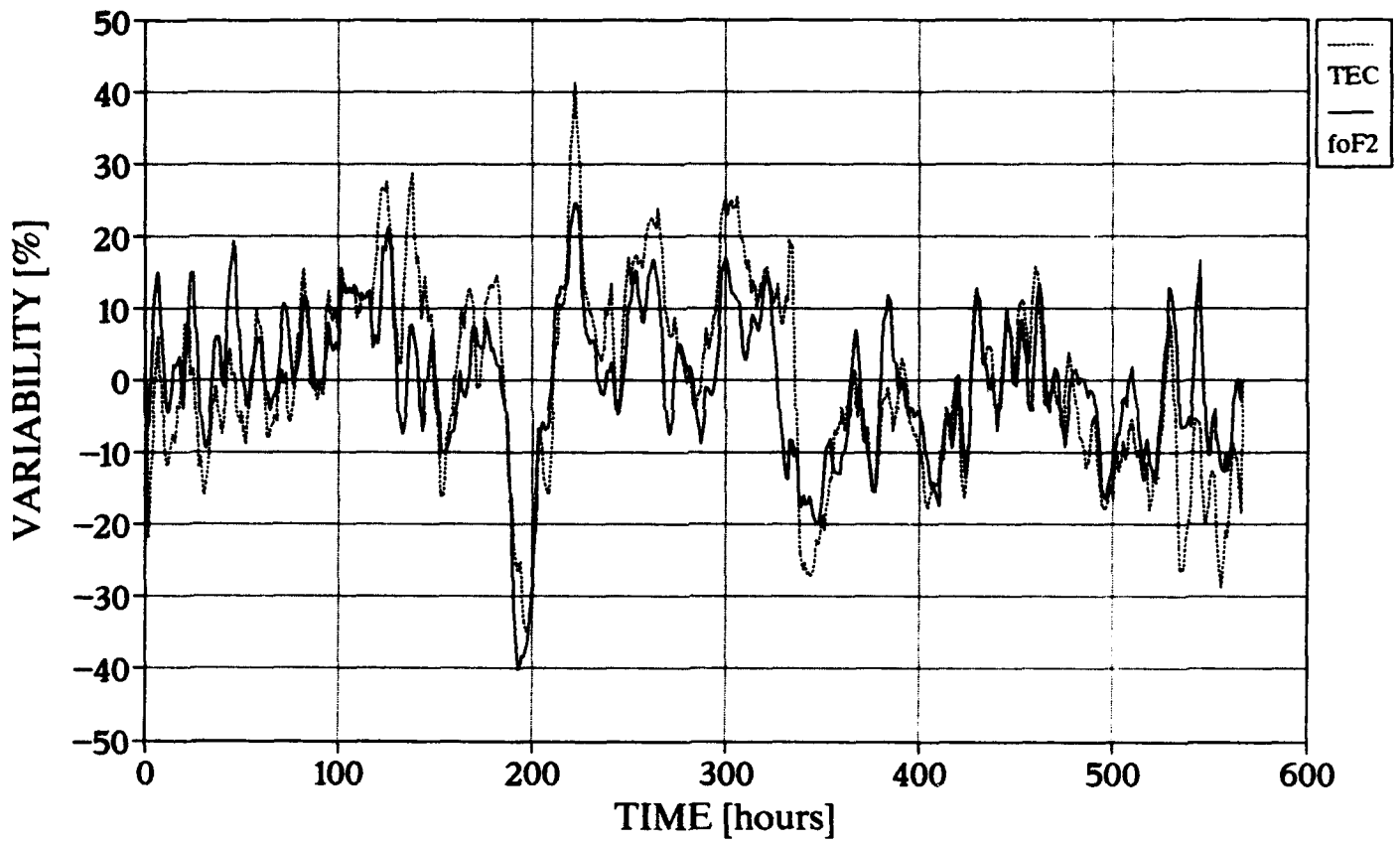


Fig B-27

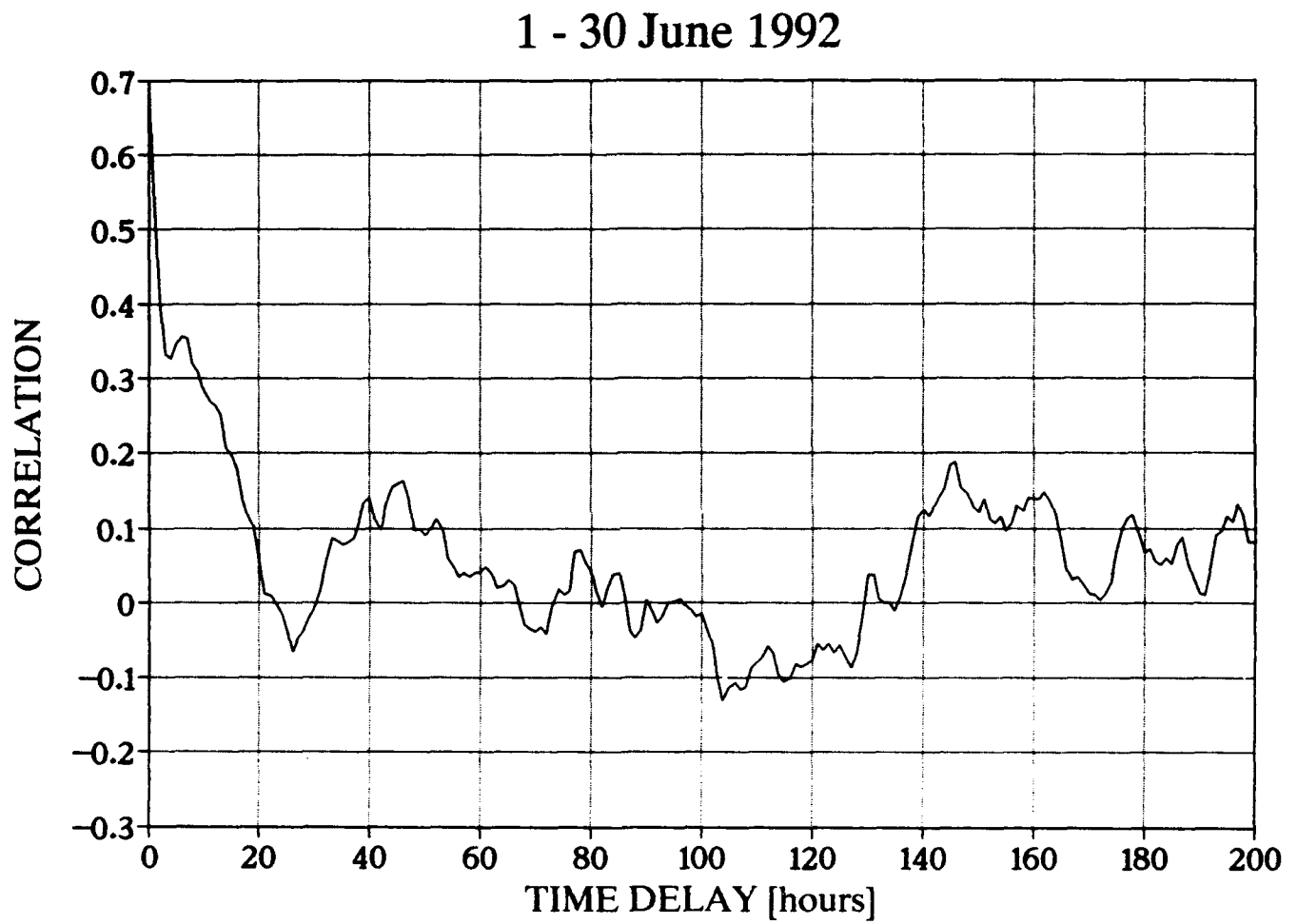


Fig B-28

1 - 30 June 1992

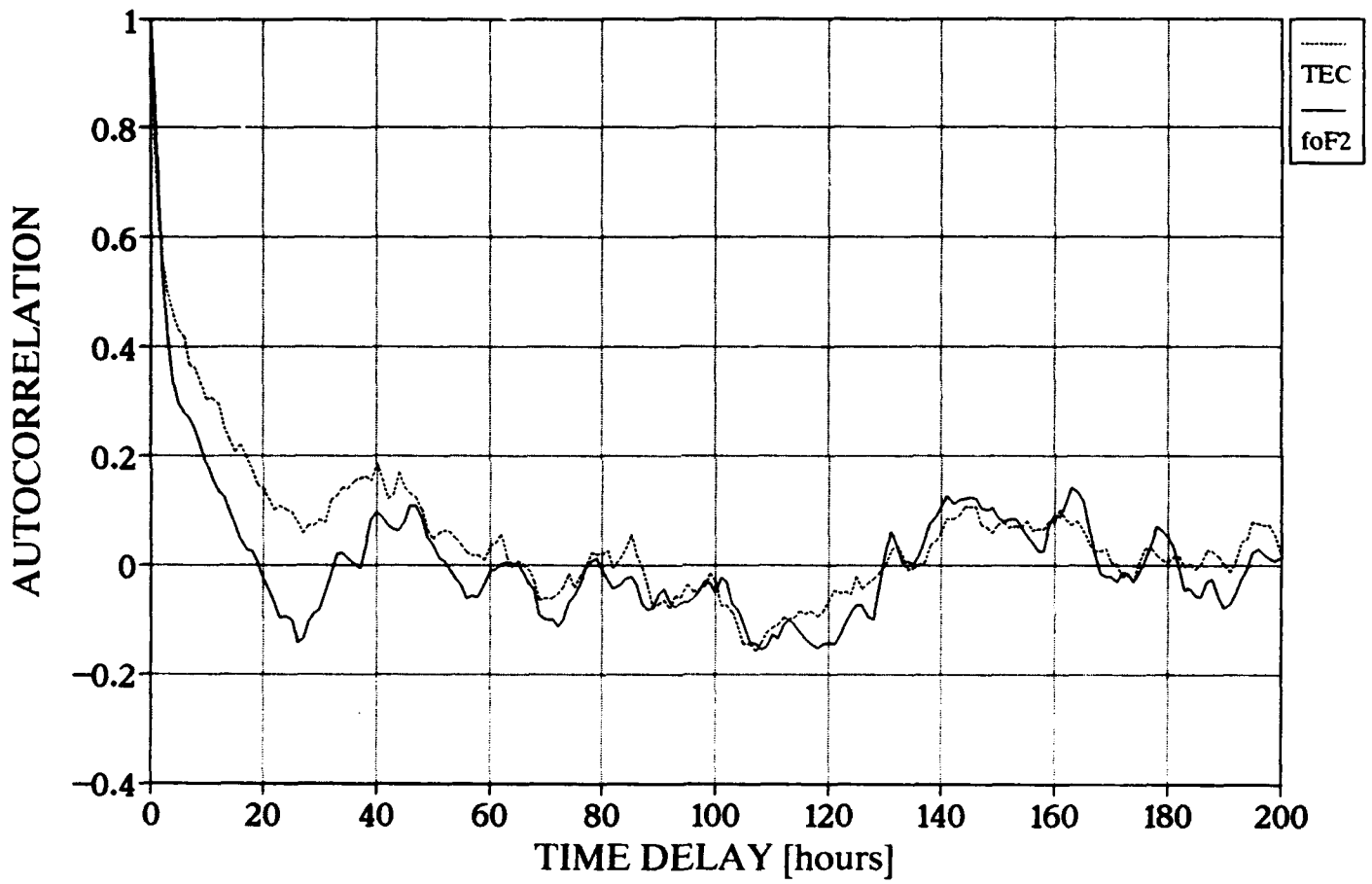


Fig B-29

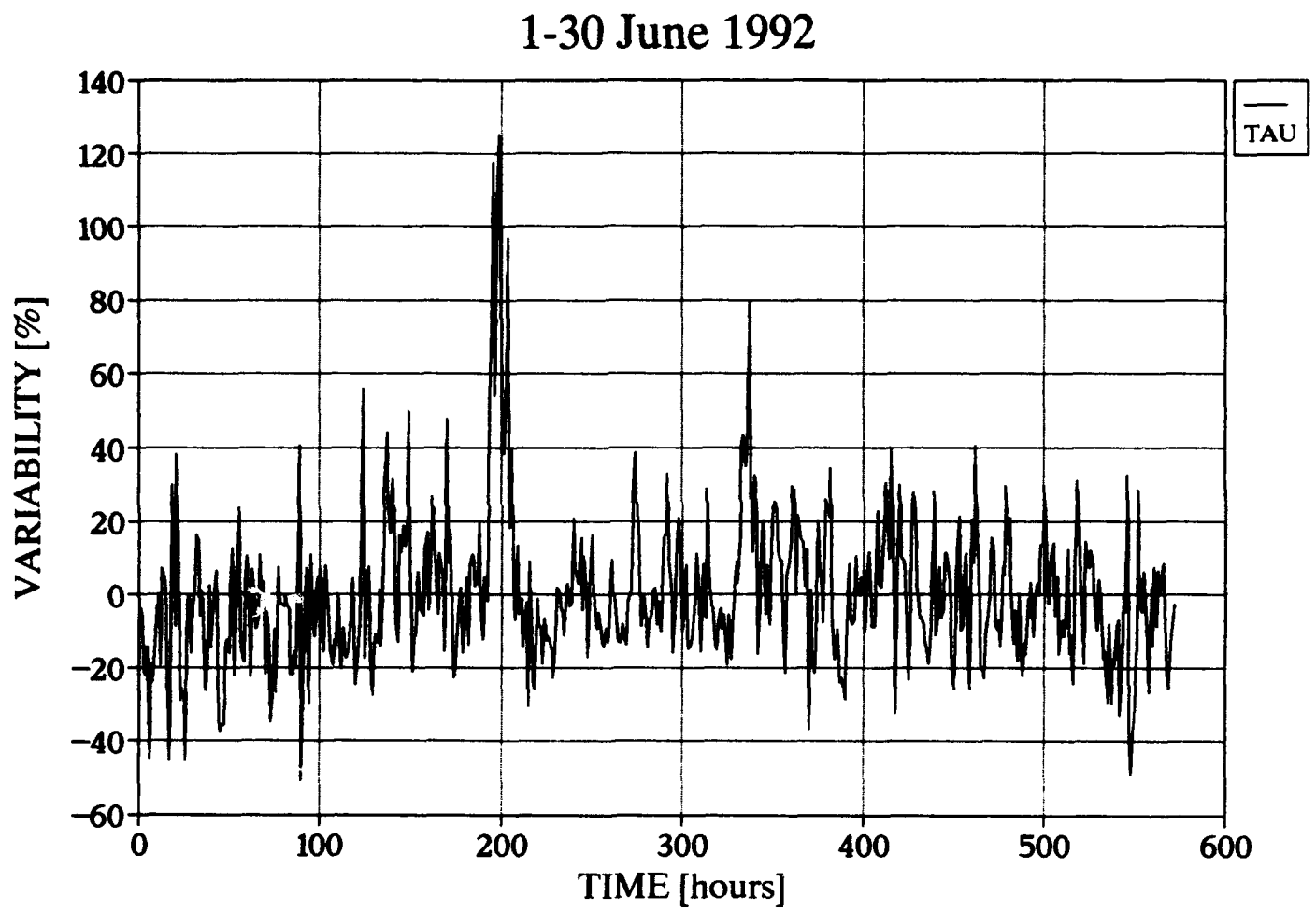


Fig B-30

2-31 July 1992
 $C=1.778$

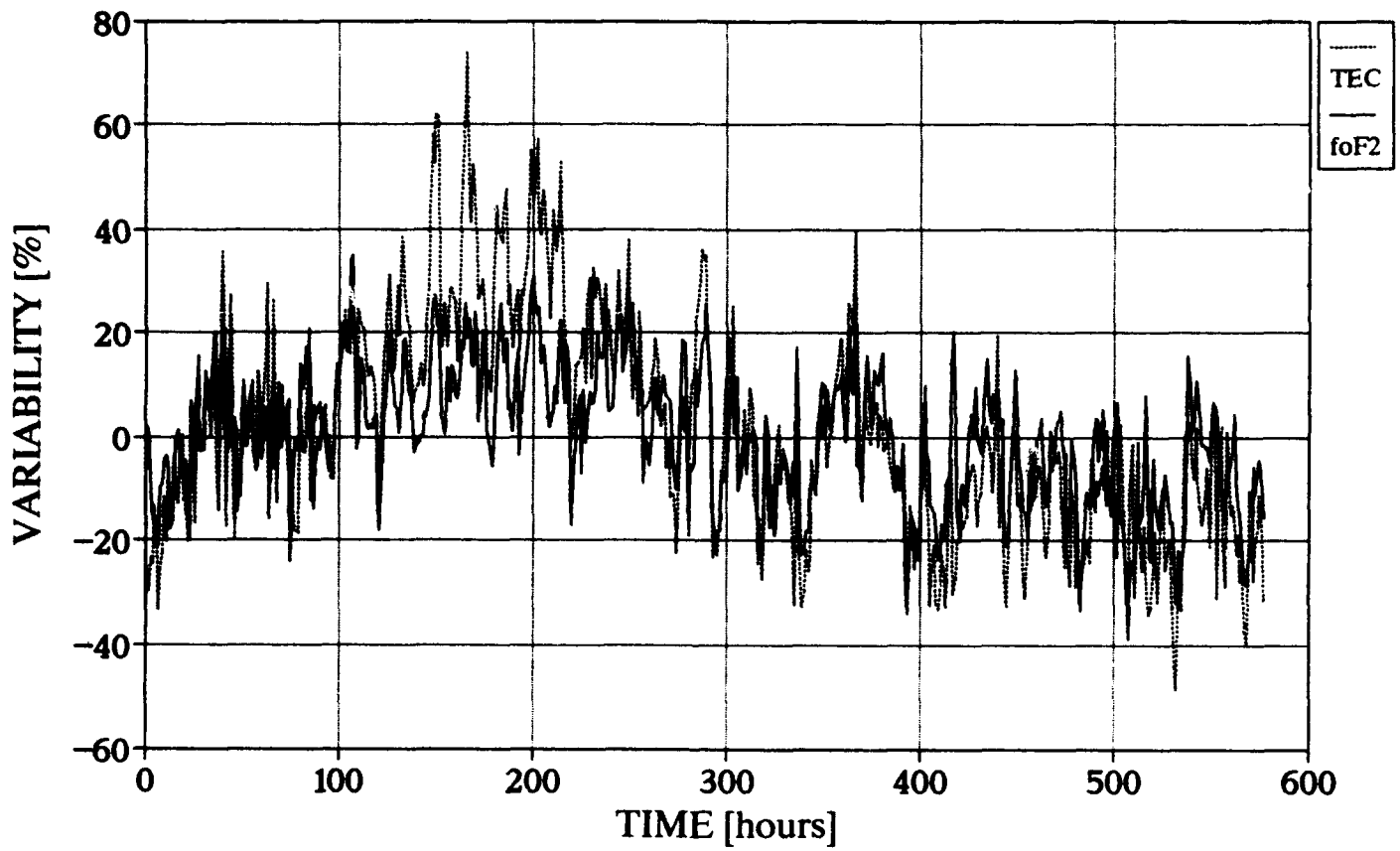


Fig B-31

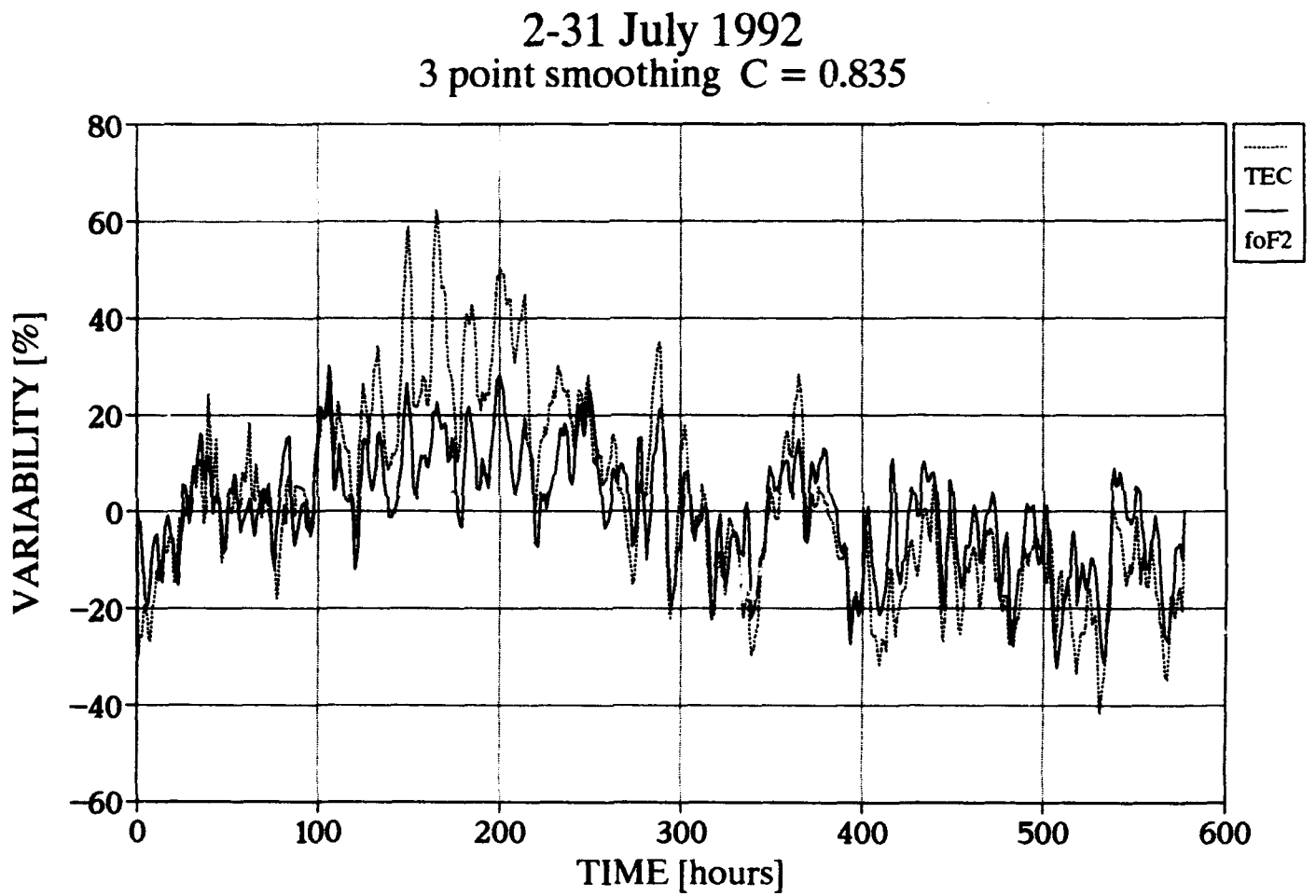


Fig B-32

2-31 July 1992
5 point smoothing $C = 0.859$

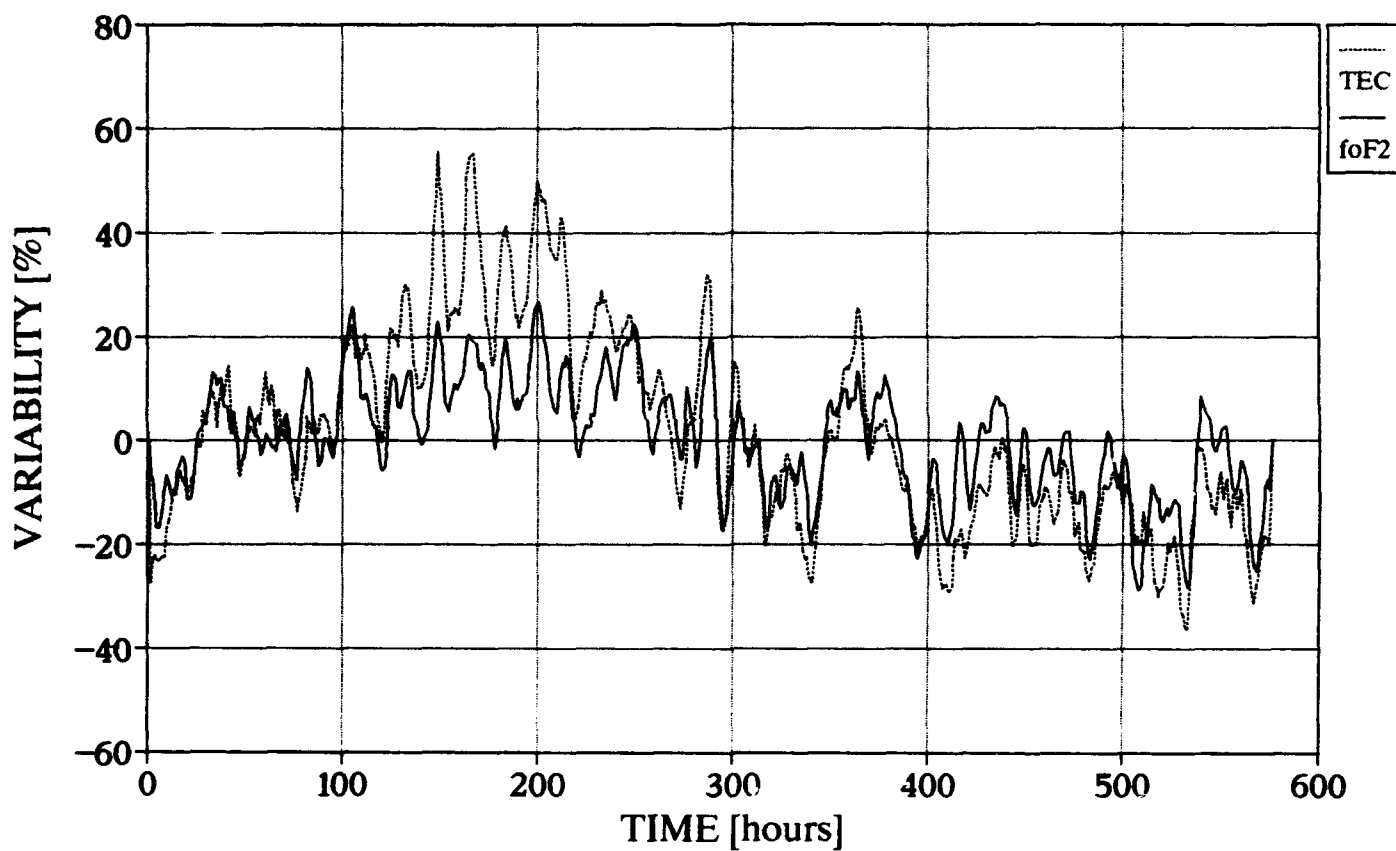


Fig B-53

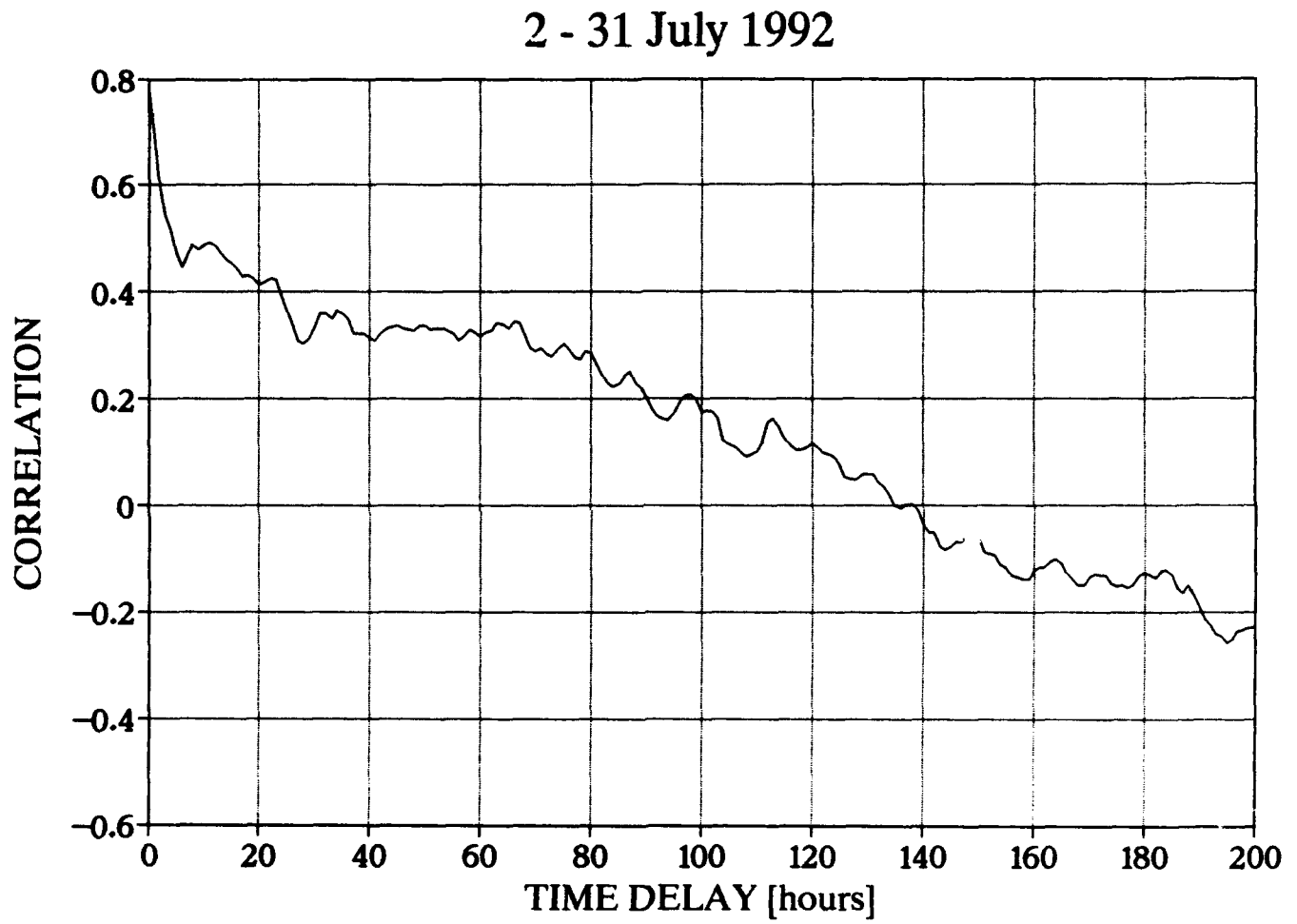


Fig B-34

2 - 31 July 1992

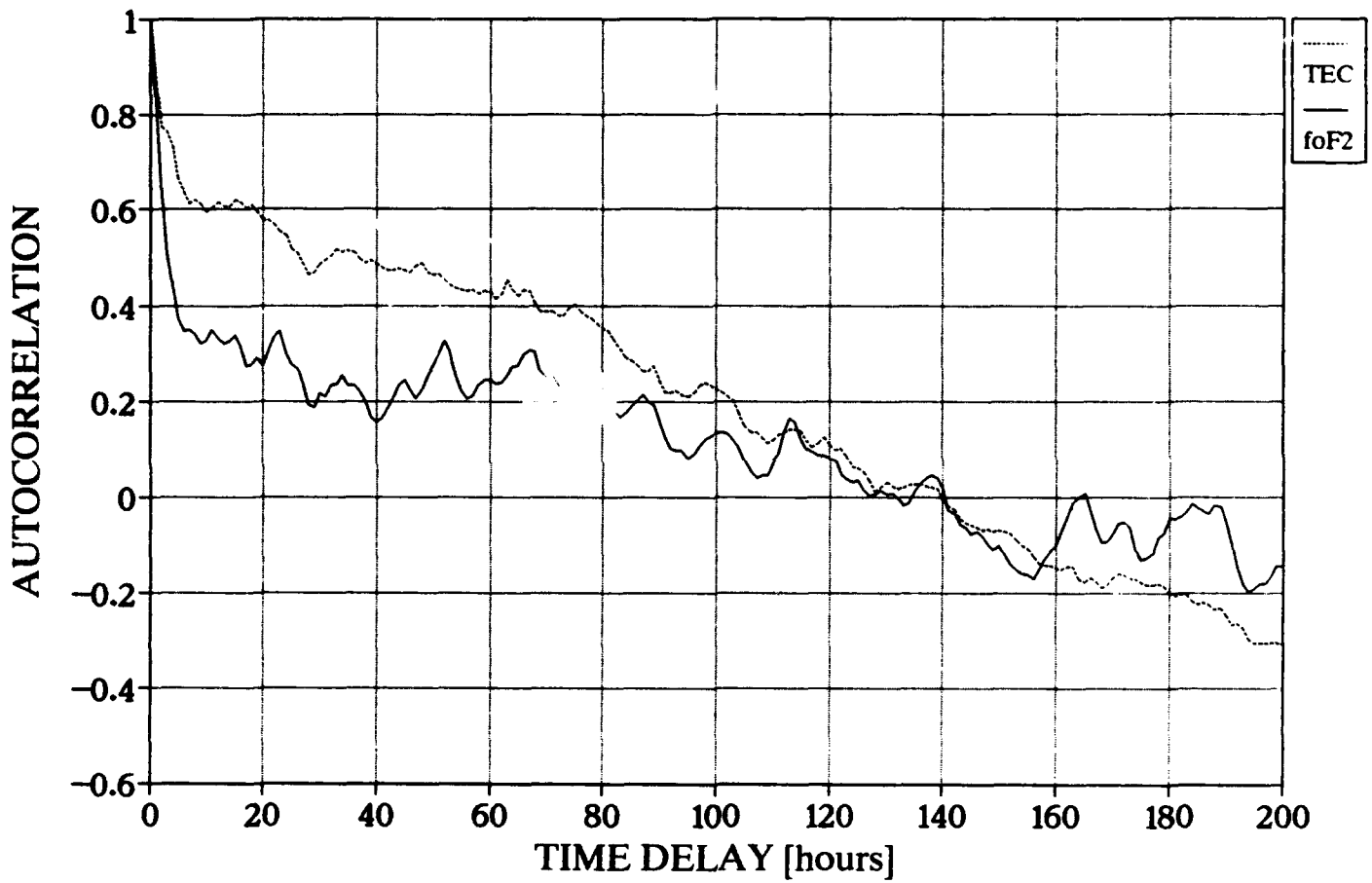


Fig B-35

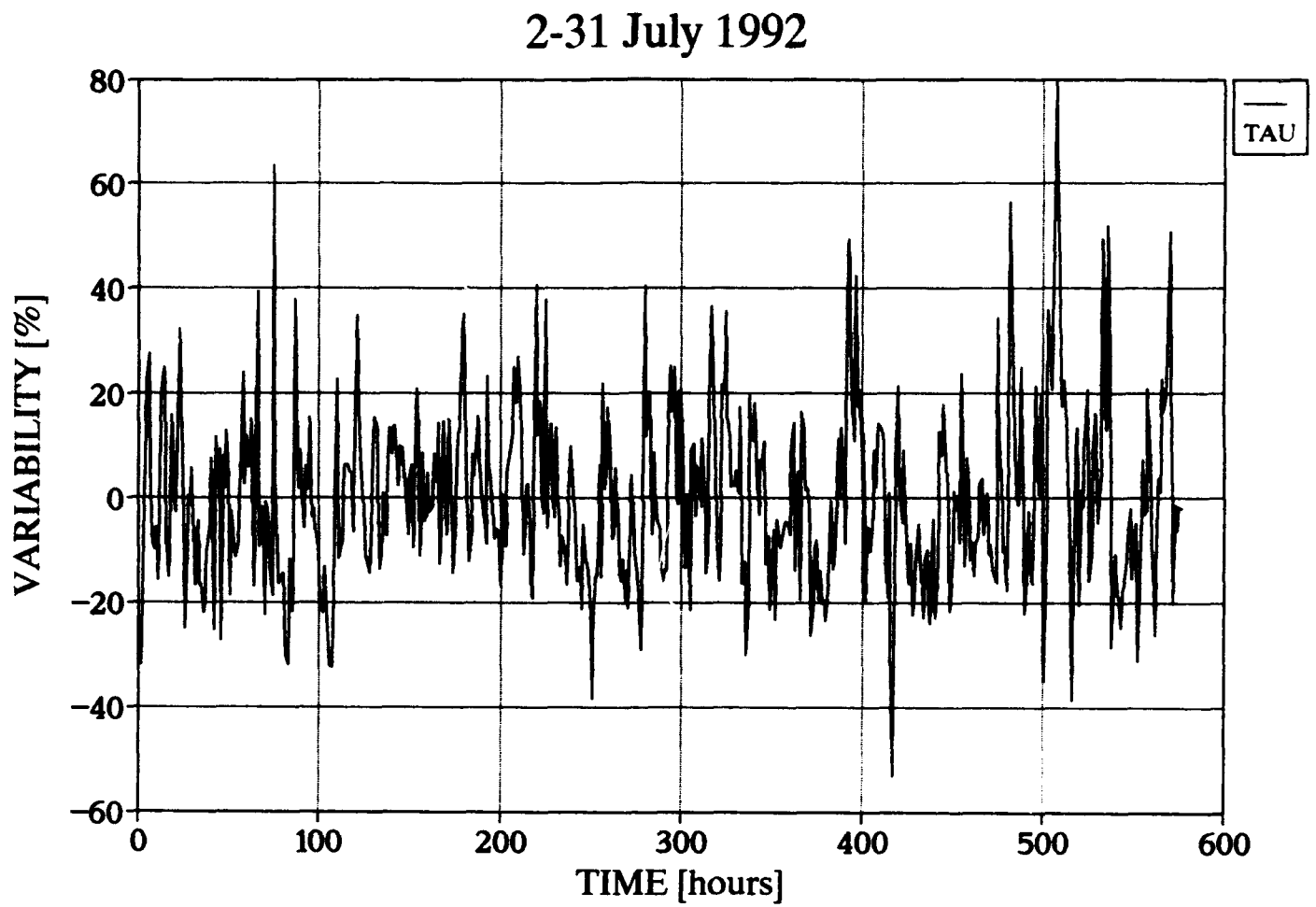


Fig B-36

1-16 August 1992
 $C=0.674$

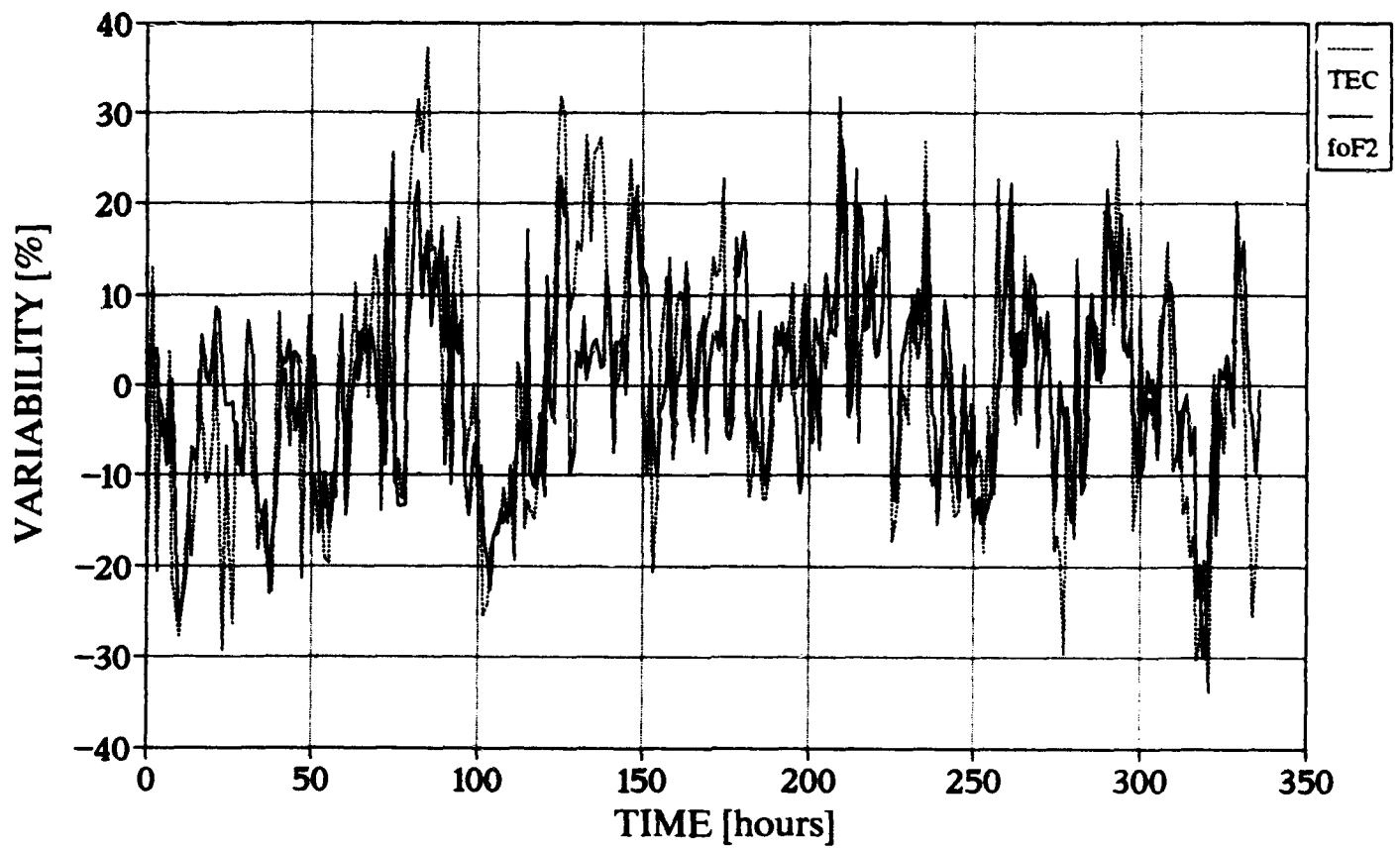


Fig B-37

1 - 16 August 1992
3 point smoothing $C = 0.792$

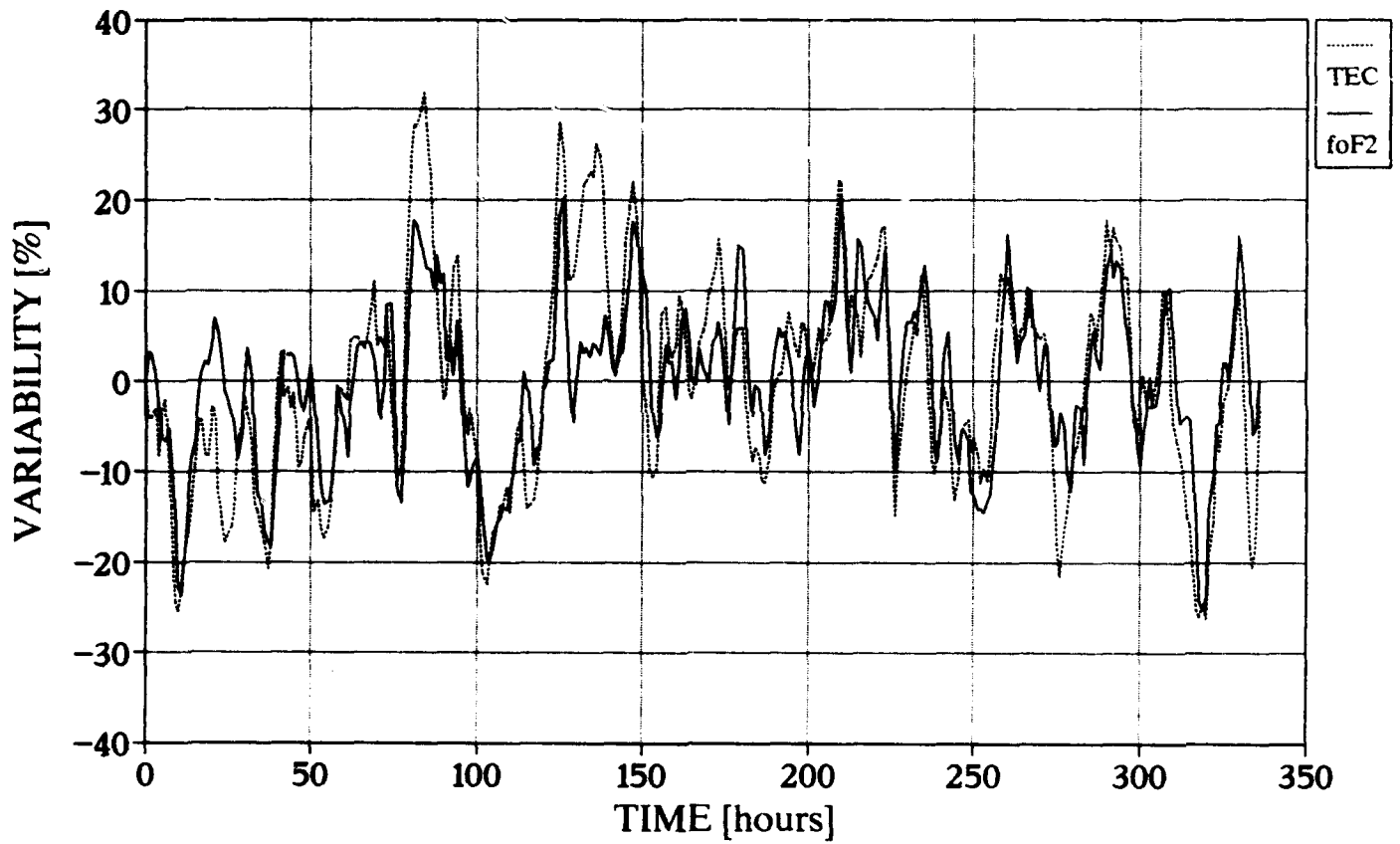


Fig B-38

1 - 16 August 1992
5 point smoothing $C = 0.823$

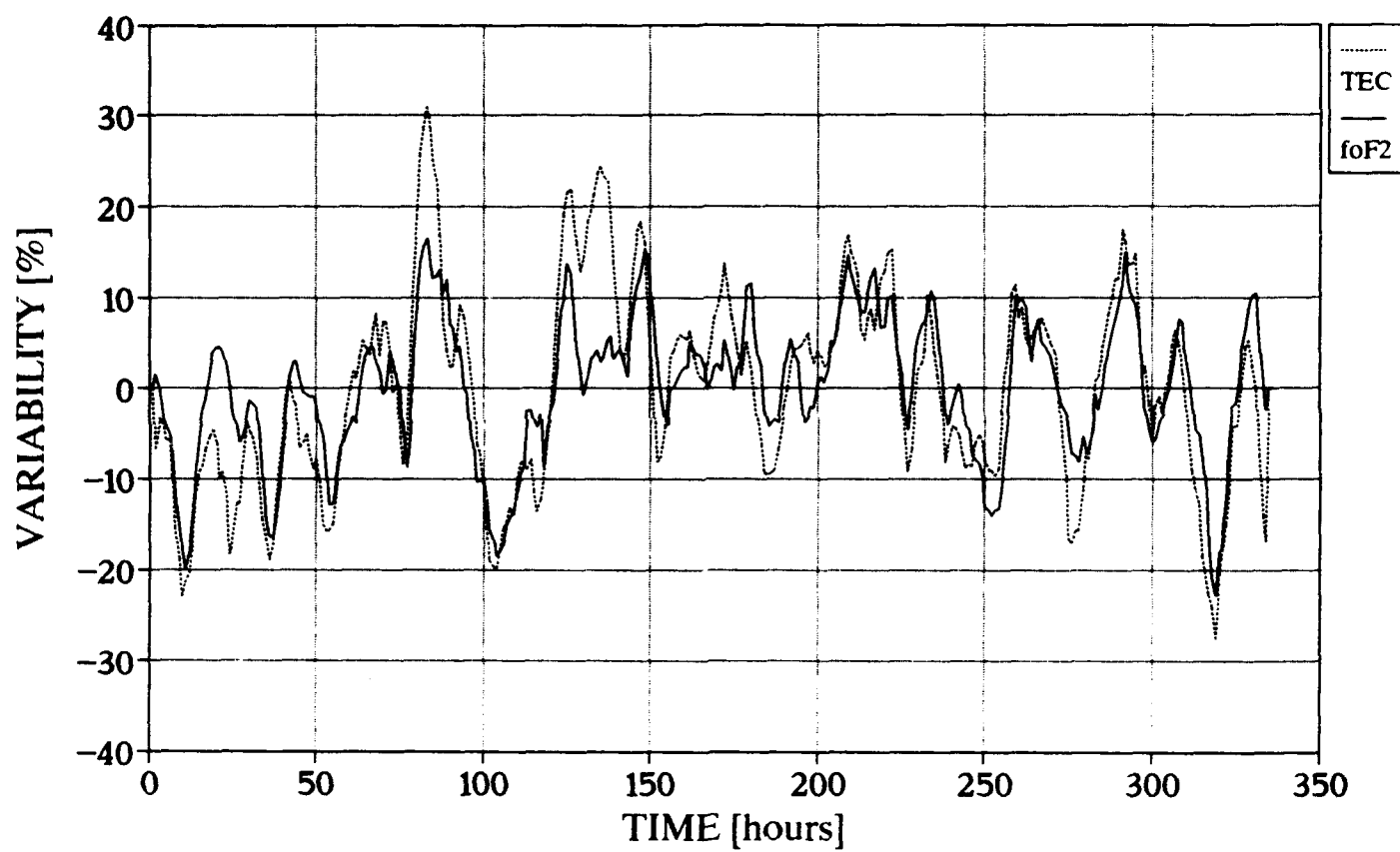


Fig B-39

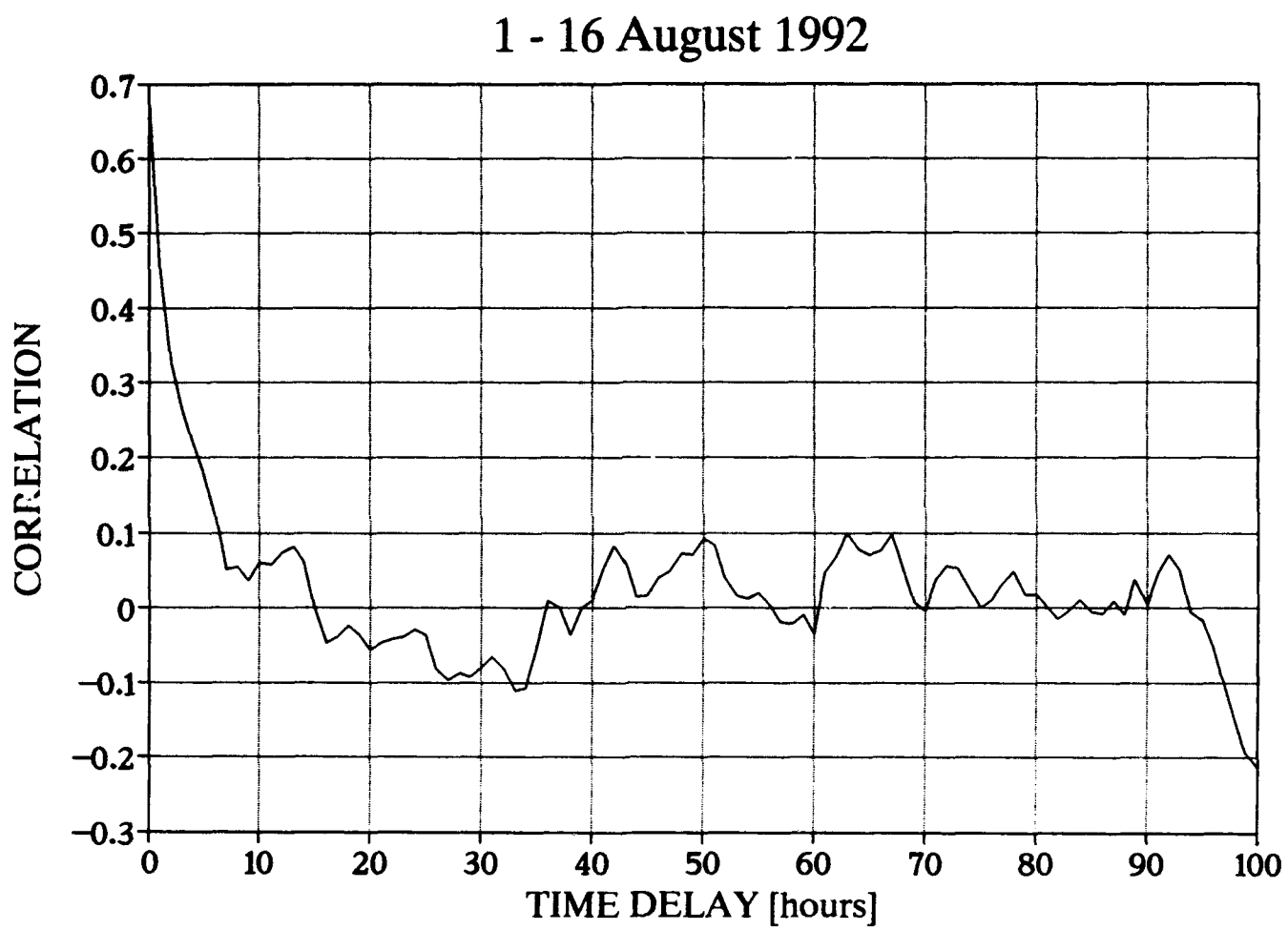


Fig B-40

1 - 16 August 1992

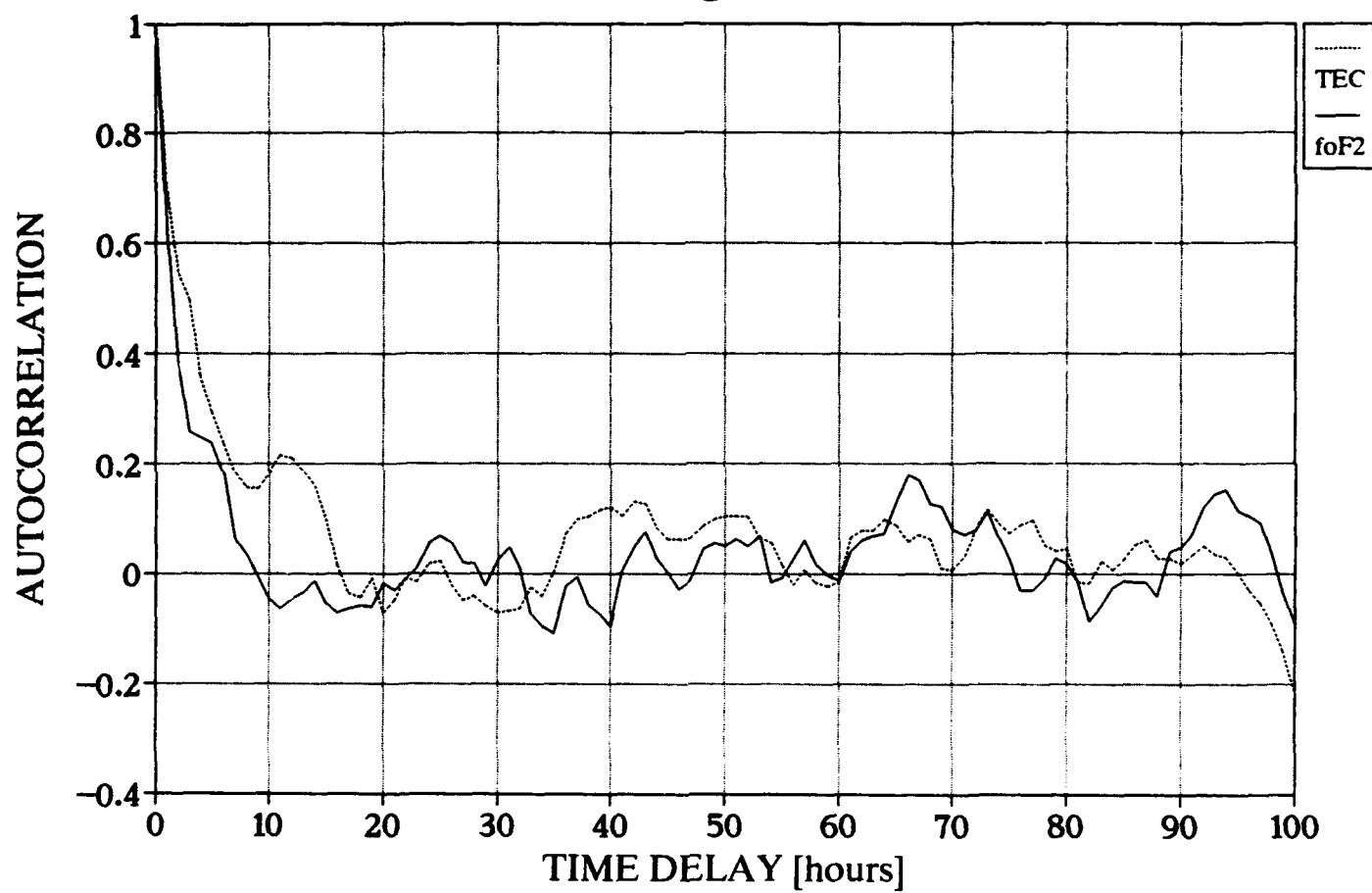


Fig B-41

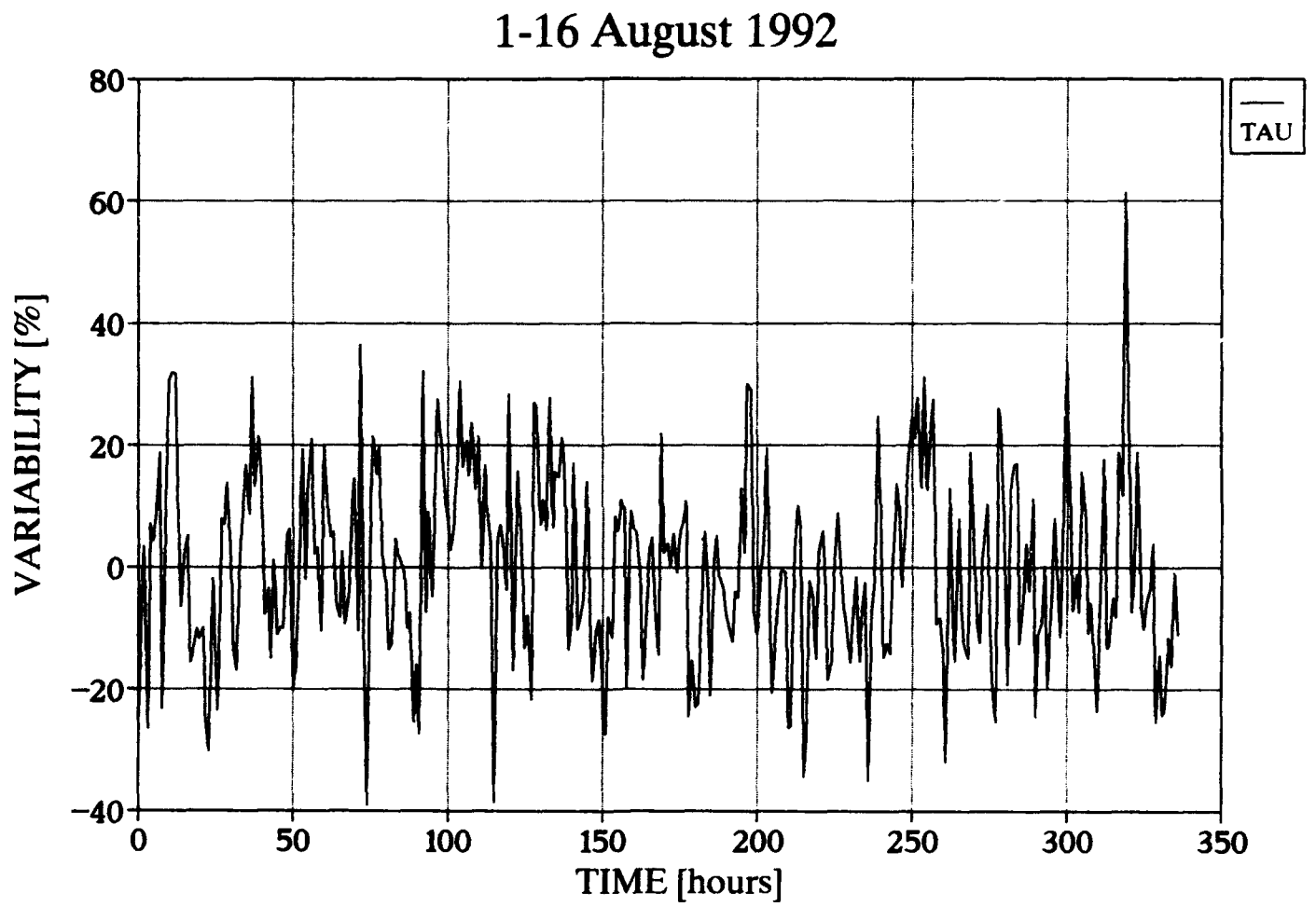


Fig B-42